

Effective Shell-Model Hamiltonians from Realistic NN Potentials

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Elba XII Workshop - Electron-Nucleus Scattering XII



A brief introduction for non-experts

What is a realistic effective shell-model hamiltonian ?



An example: ^{19}F

^{19}F



protons



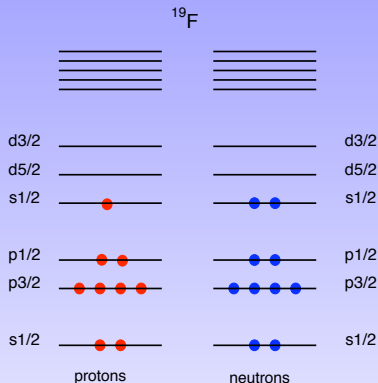
neutrons

- 9 protons & 10 neutrons interacting
- spherically symmetric mean field (e.g. harmonic oscillator)
- 1 valence proton & 2 valence neutrons interacting in a truncated model space

The degrees of freedom of the core nucleons and the excitations of the valence ones above the model space are not considered explicitly.



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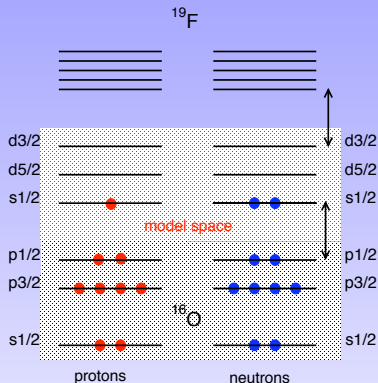


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Effective shell-model hamiltonian

The shell-model hamiltonian has to take into account in an effective way all the degrees of freedom not explicitly considered

Two alternative approaches

- phenomenological
- microscopic

$$V_{NN} (+V_{NNN}) \Rightarrow \text{many-body theory} \Rightarrow H_{\text{eff}}$$

Definition

The eigenvalues of H_{eff} belong to the set of eigenvalues of the full nuclear hamiltonian



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Workflow for a realistic shell-model calculation

- 1 Choose a realistic NN potential (NNN)
- 2 Determine the model space better tailored to study the system under investigation
- 3 Derive the effective shell-model hamiltonian by way of a many-body theory
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Realistic nucleon-nucleon potential: V_{NN}

At present several realistic V_{NNS} are available: CD-Bonn, Argonne V18, Nijmegen, ...

Strong short-range repulsion

How to handle the short-range repulsion ?

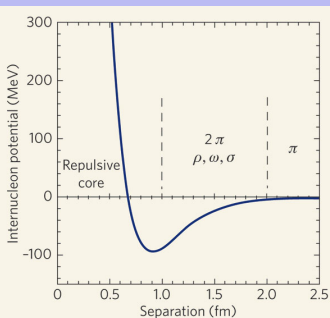
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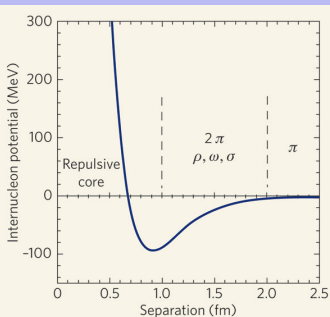
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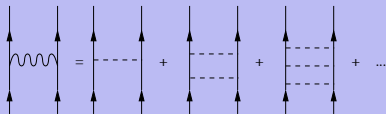
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New approach \Rightarrow low-momentum NN potentials:

- $V_{\text{low-}k}$
- chiral potentials rooted in EFT: e.g. $N^3\text{LO}$ by Entem & Machleidt (smooth cutoff $\simeq 2.5 \text{ fm}^{-1}$) or $N^3\text{LOW}$ (sharp cutoff = 2.1 fm^{-1})



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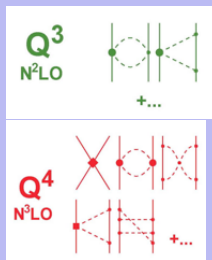
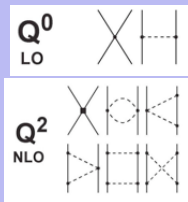


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The shell-model effective hamiltonian

A-nucleon system Schrödinger equation

$$H|\Psi_\nu\rangle = E_\nu|\Psi_\nu\rangle$$

with

$$H = H_0 + H_1 = \sum_{i=1}^A (T_i + U_i) + \sum_{i<j} (V_{ij}^{NN} - U_i)$$

Model space

$$|\Phi_i\rangle = [a_1^\dagger a_2^\dagger \dots a_n^\dagger]_i |c\rangle \Rightarrow P = \sum_{i=1}^d |\Phi_i\rangle\langle\Phi_i|$$

Model-space eigenvalue problem

$$H_{\text{eff}} P|\Psi_\alpha\rangle = E_\alpha P|\Psi_\alpha\rangle$$



The shell-model effective hamiltonian

$$\begin{pmatrix} PHP & PHQ \\ \hline QHP & QHQ \end{pmatrix} \xrightarrow{\mathcal{H} = X^{-1}HX} \begin{pmatrix} PHP & PHQ \\ \hline 0 & QHQ \end{pmatrix}$$

$QHP = 0$

$$H_{\text{eff}} = PHP$$

Suzuki & Lee $\Rightarrow X = e^{\omega}$ with $\omega = \begin{pmatrix} 0 & 0 \\ Q\omega P & 0 \end{pmatrix}$



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Folded-diagram expansion

\hat{Q} -box vertex function

$$\hat{Q}(\epsilon) = PH_1P + PH_1Q \frac{1}{\epsilon - QHQ} QH_1P$$

\Rightarrow Recursive equation for H_{eff} \Rightarrow iterative techniques
(Krenciglowa-Kuo, Lee-Suzuki, ...)

$$H_{\text{eff}} = \hat{Q} - \hat{Q}' \int \hat{Q} + \hat{Q}' \int \hat{Q} \int \hat{Q} - \hat{Q}' \int \hat{Q} \int \hat{Q} \int \hat{Q} \dots,$$

generalized folding



The perturbative approach to the shell-model H^{eff}

$$\hat{Q}(\epsilon) = PH_1P + PH_1Q \frac{1}{\epsilon - QHQ} QH_1P$$

The \hat{Q} -box can be calculated perturbatively

$$\frac{1}{\epsilon - QHQ} = \sum_{n=0}^{\infty} \frac{(QH_1Q)^n}{(\epsilon - QH_0Q)^{n+1}}$$

The diagrammatic expansion of the \hat{Q} -box



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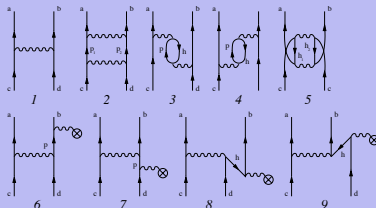
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The perturbative approach to the shell-model H^{eff}

- H^{eff} for systems with one and two valence nucleons
- \hat{Q} -box \Rightarrow Goldstone diagrams up to third order in V_{NN} (up to 2p-2h core excitations)
- Padè approximant [2|1] of the \hat{Q} -box

$$[2|1] = V_{Qbox}^0 + V_{Qbox}^1 + V_{Qbox}^2 (1 - (V_{Qbox}^2)^{-1} V_{Qbox}^3)^{-1}$$



Test case: p -shell nuclei

- $V_{NN} \Rightarrow$ chiral N^3LO potential by Entem & Machleidt
- H_{eff} for two valence nucleons outside ${}^4\text{He}$
- Single-particle energies and residual two-body interaction are derived from the theory. **No empirical input**

First, some convergence checks !



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Convergence checks

The intermediate-state space Q

Q -space is truncated: intermediate states whose unperturbed excitation energy is greater than a fixed value E_{max} are disregarded

$$|\epsilon_0 - QH_0Q| \leq E_{max} = N_{max} \hbar\omega$$

${}^6\text{Li}$ yrast states

results quite stable for
 $N_{max} \geq 20$



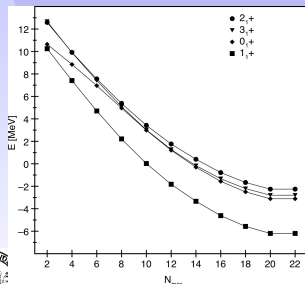
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Convergence checks

Order-by-order convergence

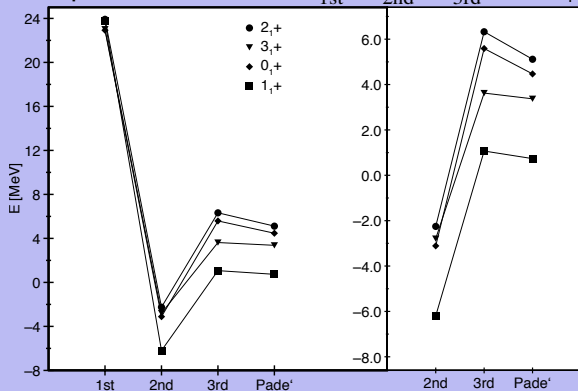
Compare results from H_{1st}^{eff} , H_{2nd}^{eff} , H_{3rd}^{eff} and $H_{Padè}^{eff}$



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Dependence on $\hbar\omega$

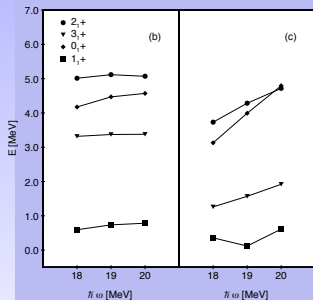
Auxiliary potential $U \Rightarrow$ harmonic oscillator potential



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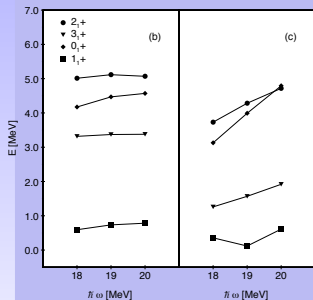
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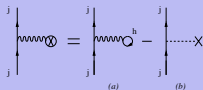
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Dependence on $\hbar\omega$

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HF-insertions



- zero in a self-consistent basis
- neglected in most applications
- neglect introduces relevant dependence on $\hbar\omega$



Benchmark calculation

Approximations are under control ... and what about the accuracy of the results ?

Compare the results with the “exact” ones

ab initio no-core shell model (NCSM)

P. Navrátil, E. Caurier, Phys. Rev. C **69**, 014311 (2004)

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Benchmark calculation

To compare our results with NCSM we need to start from a translationally invariant Hamiltonian

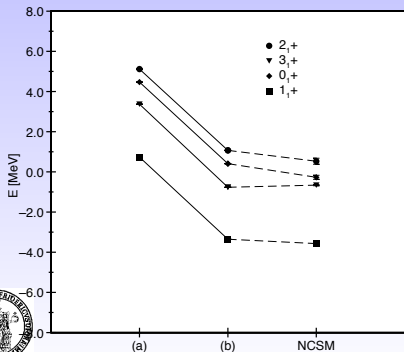
$$\begin{aligned} H_{int} &= \left(1 - \frac{1}{A}\right) \sum_{i=1}^A \frac{p_i^2}{2m} + \sum_{i<j=1}^A \left(V_{ij}^{NN} - \frac{\mathbf{p}_i \cdot \mathbf{p}_j}{mA} \right) = \\ &= \left[\sum_{i=1}^A \left(\frac{p_i^2}{2m} + U_i \right) \right] + \left[\sum_{i<j=1}^A \left(V_{ij}^{NN} - U_i - \frac{p_i^2}{2mA} - \frac{\mathbf{p}_i \cdot \mathbf{p}_j}{mA} \right) \right] \end{aligned}$$



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(a) not translationally invariant Hamiltonian
(b) purely intrinsic hamiltonian



Benchmark calculation

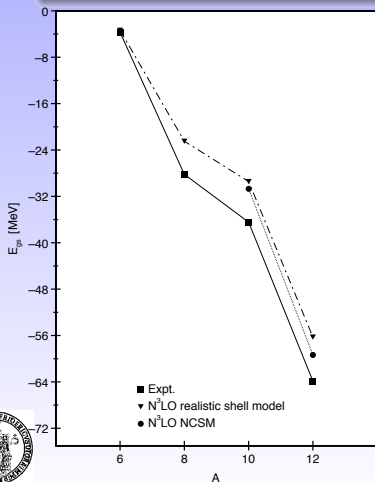
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- ground-state energies for $N = Z$ nuclei
- discrepancy grows with the number of valence nucleons

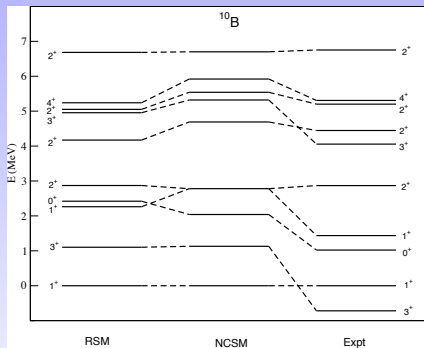


^{10}B relative spectrum



Benchmark calculation

^{10}B relative spectrum



- discrepancy ≤ 1 MeV
- minor role of many-body correlations



... At long last some physics ...

Neutron-rich Carbon isotopes

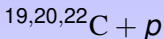
			9		F ¹⁸ 18.011 u -1.011 MeV 0- 18.9984032 2.7(2.0) ns	F ¹⁴ (2-) 0- EC	F ¹⁵ 1.0 MeV (1/2-) 0- P	F ¹⁶ 40 keV 0- EC	F ¹⁷ 64.89 s 5/2- 0- EC	F ¹⁸ 109.77 m 1- EC	F ¹⁹ 109.77 m 1/2+ 0- β	F ²⁰ 11.80 s 2+ 0- β	F ²¹ 4.158 s 5/2- 0- β	F ²² 4.23 s 4+(3+) 0- β	F ²³ 2.23 s (3/2,5/2)+ 0- β	F ²⁴ 0.36 s (1,2,3)+ 0- β	F ²⁵ 59 ms 0- β	F ²⁶ 0- β
		8	O ¹² 12.000 u 0- 15.9991 0.0719 s	O ¹³ 0.40 MeV 0- EC	O ¹⁴ 8.58 ms 0- EC	O ¹⁵ 70.696 s 0- EC	O ¹⁶ 122.23 s 1/2- 0- EC	O ¹⁷ 99.562 s 0- β	O ¹⁸ 6.038 s 0- β	O ¹⁹ 36.91 s 5/2- 0- β	O ²⁰ 13.51 s 0- β	O ²¹ 3.42 s 0- β	O ²² 2.25 s 0- β	O ²³ 82 ms 0- β	O ²⁴ 61 ms 0- β	O ²⁵ 0- β	O ²⁶ 0- β	
	7	N ¹⁰ 13.005 u 0- 15.9949 0.0051 s	N ¹¹ 760 keV 1/2- EC	N ¹² 11.600 ms 1- EC	N ¹³ 9.965 m 1/2- EC	N ¹⁴ 1- 1/2- β	N ¹⁵ 99.634 s 0- β	N ¹⁶ 71.5 s 3- β	N ¹⁷ 4.173 s 1/2- β	N ¹⁸ 628 ms 1- β	N ¹⁹ 0.304 s (1/2-) β	N ²⁰ 100 ms β	N ²¹ 85 ms β	N ²² 24 ms β	N ²³ 0- β	N ²⁴ 0- β		
6	C ⁸ 8.003 u 2.90 keV 0- 12.0007 0.0313 s	C ⁹ 126.5 ms 0- EC	C ¹⁰ 19.255 s 0- EC	C ¹¹ 20.39 m 0- EC	C ¹² 98.90 s 1/2- β	C ¹³ 1.02 s 0- β	C ¹⁴ 57.30 s 0- β	C ¹⁵ 2.449 s 1/2- β	C ¹⁶ 0.747 s 1/2- β	C ¹⁷ 193 ms β	C ¹⁸ 95 ms β	C ¹⁹ 46 ms β	C ²⁰ 14 ms β	C ²¹ 0- β	C ²² 0- β			
5	B ⁷ 7.016 u 1.4 MeV (3/2-) 0- 10.811 5.9(1.0) s	B ⁸ 770 ms 0- EC	B ⁹ 0.54 keV 3/2- 2p	B ¹⁰ 19.9 s 0- EC	B ¹¹ 80.1 s 0- β	B ¹² 20.20 ms 3/2- β	B ¹³ 17.36 ms 3/2- β	B ¹⁴ 13.4 ms 2- β	B ¹⁵ 10.5 ms 0- β	B ¹⁶ 5.00 ms 0- β	B ¹⁷ 5.08 ms (3/2-) β	B ¹⁸ 0- β	B ¹⁹ 0- β					
3e	Be ⁵ 9.012 82 3.9(1.0) s	Be ⁶ 92 keV 0- 2p	Be ⁷ 53.12 d 3/2- EC	Be ⁸ 6.8 s 0- EC	Be ⁹ 1.51E-6 s 1/2- EC	Be ¹⁰ 1381 s 1/2- β	Be ¹¹ 1381 s 1/2- β	Be ¹² 23.6 ms 0- β	Be ¹³ 0.9 MeV (1/2,3/2)+ β	Be ¹⁴ 4.38 ms 0- β								
4	Li ⁴ 4.002 603 8.75(1.0) s	Li ⁵ 1.5 MeV 3/2- 0- 6.981 8.7(1.0) s	Li ⁶ 7.5 s 0- β	Li ⁷ 92.5 s 0- β	Li ⁸ 838 ms 2- β	Li ⁹ 178(3) ms 3/2- β	Li ¹⁰ 1.2 MeV 3/2- β	Li ¹¹ 8.5 ms 3/2- β	Li ¹² 0- β									
3e	He ³ 3.016 049 8.375 s	He ⁴ 4.002 603 0- EC	He ⁵ 0.60 MeV 3/2- β	He ⁶ 86(7) ms 0- β	He ⁷ 160(2) ms 0- β	He ⁸ 115(6) ms 1/2- β	He ⁹ 1.2 MeV (1/2-) β	He ¹⁰ 0.3 MeV 0- β										





Observation of a Large Reaction Cross Section in the Drip-Line Nucleus ^{22}C

K. Tanaka,¹ T. Yamaguchi,² T. Suzuki,² T. Ohtsubo,³ M. Fukuda,⁴ D. Nishimura,⁴ M. Takechi,^{4,1} K. Ogata,⁵ A. Ozawa,⁶
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K. Yamada,¹ T. Yasuno,⁶ and M. Yoshitake²



10 ^{22}C counts per hour

TABLE I. Reaction cross sections (σ_R) in millibarns.

A	σ_R
19	754(22)
20	791(34)
22	1338(274)

RIKEN Radioactive Isotope Beam Factory (RIBF)



^{22}C is the heaviest Borromean nucleus ever observed

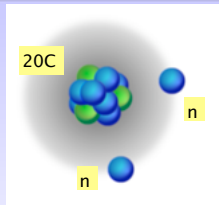


Borromean nucleus \Rightarrow weakly bound nucleus that has, considered as a three-body system, no bound states in the binary subsystems.

^{22}C is weakly bound $S_{2n}^{\text{est}} = 420\text{keV}$

^{21}C is unstable for one-neutron emission

M. Langevin *et al.*, Phys. Lett. B **150**, 71 (1985)



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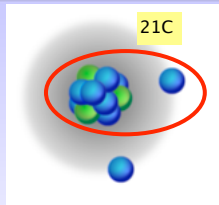


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PHYSICAL REVIEW C **81**, 064303 (2010)

Shell-model calculations for neutron-rich carbon isotopes with a chiral nucleon-nucleon potential

L. Coraggio,¹ A. Covello,^{1,2} A. Gargano,¹ and N. Itaco^{1,2}

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²*Dipartimento di Scienze Fisiche, Università di Napoli Federico II, Complesso Universitario di Monte S. Angelo,
Via Cintia, I-80126 Napoli, Italy*

- Realistic potential $V_{NN} \Rightarrow N^3\text{LOW}$
- $\Rightarrow H_{\text{eff}}$
- Single-particle energies & effective interaction from the theory. **No empirical input**



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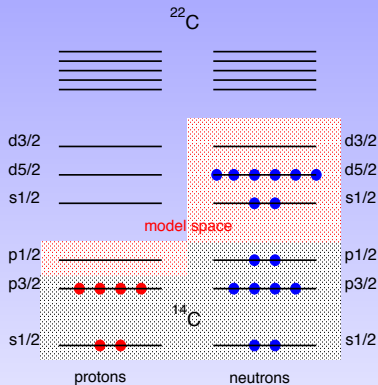
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Model space



^{14}C inert core

Neutrons:

shell $sd \Rightarrow 1s1/2, 0d5/2, 0d3/2$

Protons:

$0p1/2$



- ^{22}C & drip line
- $N=14$ shell closure

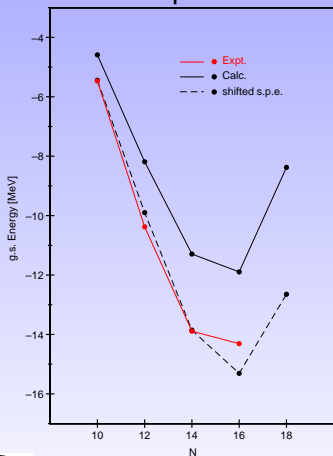
Single-particle energies

nlj	ϵ MeV
$1s1/2$	0.0 (-0.793)
$0d5/2$	5.914
$0d3/2$	1.394



Results: ^{22}C & drip line

g.s energies of even mass
Carbon isotopes



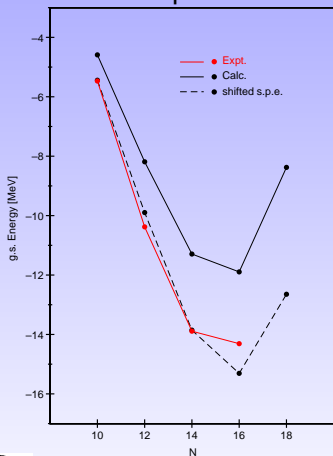
$$S_n(^{21}\text{C}) = -1.6 \text{ MeV}$$

$$S_{2n}(^{22}\text{C}) = 456 \text{ keV}$$



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Results: N=14 shell closure

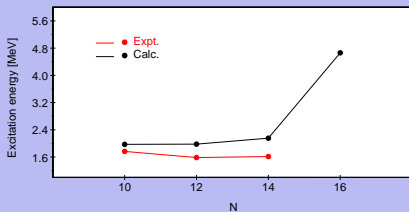
Experimentally the N=14 shell closure in oxygen isotopes disappears in carbon isotopes



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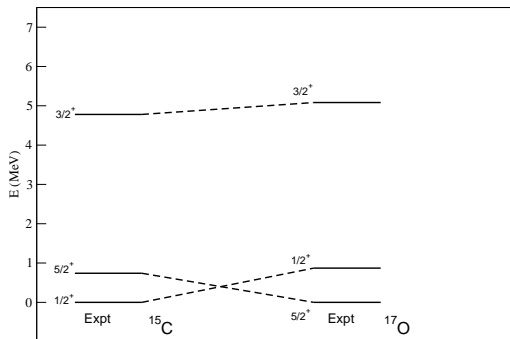
Excitation energies of 2_1^+ states in heavy carbon isotopes



Results: N=14 shell closure

The pn monopole component of V_{eff} reproduces correctly the g.s. inversion in ^{15}C and ^{17}O

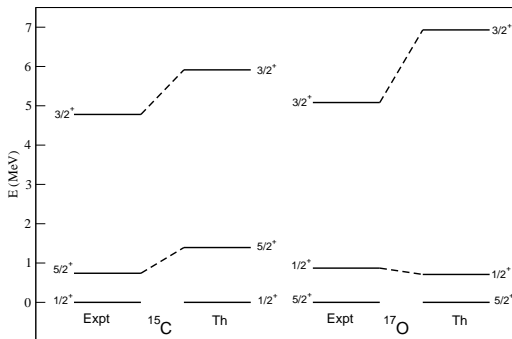
$$\epsilon_j(^{17}\text{O}) = \epsilon_j(^{15}\text{C}) + V_{p_{2j}^1}^{\text{mon}} \times 2$$



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$$\epsilon_j(^{17}\text{O}) = \epsilon_j(^{15}\text{C}) + V_{p\frac{1}{2};j}^{\text{mon}} \times 2$$



Concluding remarks

- Approximations involved may be kept under control
- Effective shell-model hamiltonians derived from realistic NN potentials are reliable
- Realistic shell-model calculations \Rightarrow fully microscopic description of nuclear structure
 - no parameters modified *ad hoc*
 - enhanced predictive power
- Effective realistic shell-model hamiltonians may be employed successfully to describe the properties of nuclei close to the drip lines
 - drip-lines position
 - properties of loosely bound nuclei
 - evolution of nuclear properties when varying N/Z
 - shell-quenching
 - new magic numbers



Effective Shell-Model Hamiltonians from Realistic NN Potentials

Nunzio Itaco

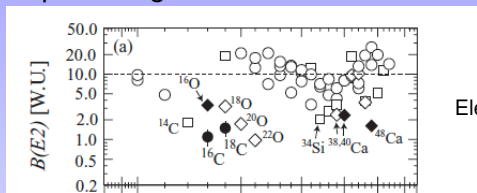
Università di Napoli Federico II
Istituto Nazionale di Fisica Nucleare - Sezione di Napoli

Elba XII Workshop - Electron-Nucleus Scattering XII



Results: $B(E2; 2_1^+ \rightarrow 0_1^+)$ transition probabilities

Reduced Experimental $B(E2; 2_1^+ \rightarrow 0_1^+) \Rightarrow$ effective charges quenching



Ong *et al* Phys. Rev. C **78**, 014308 (2008)

Elekes *et al* Phys. Rev. C **79**, 011302 (2009)

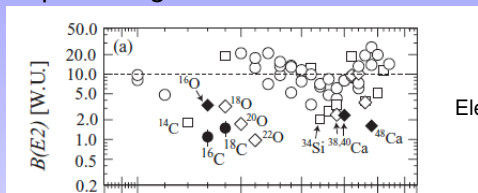
$B(E2; 2_1^+ \rightarrow 0_1^+)$ in $e^2 fm^4$

	Calc.	Expt.
^{16}C	1.8	$2.6 \pm 0.2 \pm 0.7$ 4.15 ± 0.73
^{18}C	3.0	$4.3 \pm 0.2 \pm 1.0$
^{20}C	3.7	< 3.7



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