# Nuclear effects on the determination of neutrino oscillation parameters

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Enrique Fernandez Martinez and Jose Udias



## Main motivation of this work

comparing Fermi gas (FG) and advanced nuclear model predictions for physically interesting neutrino observables

- this is relevant because many MonteCarlo codes, used to study the sensitivity to still *unknown* parameters at future  $\nu$  facilities are based on FG
- impossible to discuss all recent nuclear models
- focus the attention on two different approaches

- Introduction
  - The Standard Model of neutrino oscillations
  - What we know and what we do not know
  - The importance of  $\theta_{13}$  and  $\delta$  !
- The nuclear cross sections
  - Nuclear cross sections in the QE region
  - The QE region
  - The Spectral Function Approach
  - The Relativistic Mean Field approximation
  - The Relativistic Fermi Gas Model
  - The  $\nu$ -nucleus cross sections
- Facility and observables
  - The βBeam facility
  - The CP discovery potential
  - The sensitivity to  $\theta_{13}$
  - A combined analysis
  - Generalizing the previous results
- Summary and conclusions



The Standard Model of neutrino oscillations

# ν FLAVOUR CONVERSION has been confirmed in many experiments

$$U = R_{23}(\theta_{23})R_{13}(\theta_{13}, \delta)R_{12}(\theta_{12})$$

The neutrino oscillation probability (in matter)

$$P_{\alpha\beta} = \left| A_{\alpha\beta} \right|^2 = \sum_{i,j} \tilde{U}_{\alpha i}^* \tilde{U}_{\beta i} \tilde{U}_{\alpha j} \tilde{U}_{\beta j}^* \exp\left(i \frac{\tilde{m}_j^2 - \tilde{m}_i^2}{2E} L\right)$$

E is the neutrino energy, L is the baseline length,  $\tilde{m}_i$  and  $\tilde{U}_{\beta j}$  are the mass of the ith neutrino mass eigenstate and the mixing matrix in matter

- Usual assumption: U is a  $3 \times 3$  unitary mixing matrix
- lacktriangle three angles  $heta_{ij}$  and one CP phase  $\delta$



the standard framework implies 7 parameters to describe  $\nu$  oscillation in matter\_



## Global 3 $\nu$ fit to the world neutrino data

## At $1\sigma$ (3 $\sigma$ )

M. C. Gonzalez-Garcia and M. Maltoni, Phys. Rept. 460, 1 (2008)

## well known parameters

$$\begin{split} \Delta m_{21}^2 &= 7.67 ^{+0.22}_{-0.21} \left(^{+0.67}_{-0.61}\right) \times 10^{-5} \; eV^2 \;, \\ \Delta m_{31}^2 &= \begin{cases} -2.37 \pm 0.15 \; \left(^{+0.43}_{-0.46}\right) \times 10^{-3} \; eV^2 & \text{(inverted hierarchy)} \;, \\ +2.46 \pm 0.15 \; \left(^{+0.47}_{-0.42}\right) \times 10^{-3} \; eV^2 & \text{(normal hierarchy)} \;, \end{cases} \\ \theta_{12} &= 34.5 \pm 1.4 \; \left(^{+4.8}_{-4.0}\right) \;, \\ \theta_{23} &= 42.3 ^{+5.1}_{-3.3} \; \left(^{+11.3}_{-7.7}\right) \;, \end{split}$$

## poor and unknown parameters

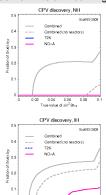
$$\theta_{13} = 0.0^{+7.9}_{-0.0} {+12.9 \choose -0.0}$$
 recent claim:  $\sin^2 \theta_{13} = 0.016 \pm 0.01$  at  $1 \sigma$ 
 $\delta_{\text{CP}} \in [0, 360] \text{ (unknown)}$  G. L. Fogli et at., arXiv:  $0806.2649$ 
 $\sin(\Delta m_{31}^2)$  octant of  $\theta_{23}$  Majorana or Dirac Neutrinos?

The importance of  $\theta_{13}$  and  $\delta$ !

# Great interest on $\theta_{13}$ and $\delta$

Introduction

## some hints at incoming experiments?



modified from P. Huber et al. JHEP 0911:044,2009

Many future experiments will look for a precise measurement of  $\theta_{13}$ .

In the standard parametrization, large  $\theta_{13}$  means good chance to reveal the CP violation in the leptonic sector

## One needs to control:

- flux composition
- detector response
- nuclear cross sections

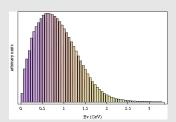


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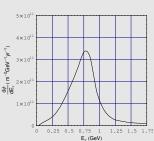
# The importance of cross sections in the QE region

many current and planned experiments use a  $\nu$  flux at energies  $\leq$  1 GeV

## MiniBoone



T2K-I



and many others (NO $\nu$ A, high  $\gamma$   $\beta$ -beams...)

very few neutrino scattering data

important to estimate precisely the  $\nu$ -nucleus cross sections in the QE region

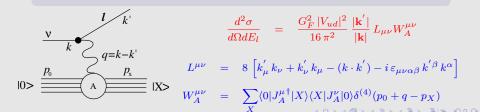


# The QE region

- at low energies ( $E_{\nu} \leq 0.6 0.7$  GeV): the dominant contribution comes from **quasi-elastic** scattering;
- at higher energies: **inelastic production** of charged leptons (via resonance excitation) + inelastic production of  $\pi^0$  also contribute
- lacksquare negligible deep inelastic scattering contribution at  $\mathcal{O}(1)$  GeV

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## formalism to describe inclusive $\nu + A \rightarrow l + X$ reaction



The QE region

# The Impulse Approximation

# the problem is the calculation of the hadronic tensor $W^{\mu\nu}_{_{A}}$

- $\mathbf{p}$  for  $|\mathbf{q}| < 0.5 \,\mathrm{GeV}$  NMBT + nonrelativistic wave functions + expansion of the current operator in powers of  $|\mathbf{q}|/m_N$  carlson&Schiavilla, Rev. Mod. Phys. 70, 743 (1998)
- for larger |q| (the energy regime we are interested in) we can no longer describe the final states  $|X\rangle$  in terms of nonrelativistic nucleons

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we need a set of simplifying assumptions to describe relativistic motion of final state particles and the occurrence of inelastic processes

the Impulse Approximation

target nucleus seen as a collection of individual nucleons  $J_{\mu} \rightarrow \sum_{i} j_{\mu}^{i}$ 

but see the Ankowsky's talk and Ankowsky et al., 1001,0481

scattered nucleons and recoiling system  ${\cal R}$  evolve independently of one another

$$|X
angle 
ightarrow |i,p^{'}
angle \otimes |\mathcal{R},p_{\mathcal{R}}
angle$$
 (no Final State Interactions)

# The Spectral Function Approach

Benhar et al., Phys.Rev.D72:053005.2005

$$\sigma \sim \sum_{i} \left| \begin{array}{c} \frac{1}{k^{i}} \\ \frac{q=k-k^{i}}{p} \\ \frac{p}{i} \end{array} \right|$$

$$\boldsymbol{\sigma} \sim \boldsymbol{\Sigma}_{i} \begin{bmatrix} \frac{\mathbf{v}}{\mathbf{k}} & \frac{\mathbf{d}^{2}\sigma_{IA}}{\mathbf{d}^{2}\mathbf{d}E_{l}} & = \int d^{3}p \, dE \, P(\mathbf{p}, E) \, \frac{d^{2}\sigma_{\text{elem}}}{d\Omega dE_{l}} \\ \frac{\mathbf{d}^{2}\sigma_{\text{elem}}}{\mathbf{d}^{2}\sigma_{\text{elem}}} & = \frac{G_{F}^{2} \, V_{ud}^{2}}{32 \, \pi^{2}} \, \frac{|\mathbf{k}'|}{|\mathbf{k}|} \, \frac{1}{4 \, E_{\mathbf{p}} \, E_{|\mathbf{p}+\mathbf{q}|}} \, L_{\mu\nu} W_{A}^{\mu\nu} \end{bmatrix}$$

$$W_A^{\mu\nu} = \frac{1}{2} \int d^3p \, dE \, P(\mathbf{p}, E) \frac{1}{4 \, E_{\mathbf{p}} \, E_{|\mathbf{p}+\mathbf{q}|}} W^{\mu\nu}(\tilde{p}, \tilde{q})$$

 $ightharpoonup P(\mathbf{p},E)$  is the target spectral function: probability distribution of finding a nucleon with momentum  $\mathbf{p}$  and removal energy E in the target nucleus

it encodes all the informations about the initial struck particle

# The Spectral Function

A. Ramos, A. Polls, W. H. Dickhoff, Nucl., Phys. A503, (1989) 1 O. Benhar, A. Fabrocini, S. Fantoni, Nucl., Phys. A505, (1989) 267 O. Benhar, A. Fabrocini, S. Fantoni and I. Sick, Nucl., Phys. A579, (1994) 493

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- the calculation of  $P(\mathbf{p}, E)$  for any A is a complicated task
- for nuclei from Carbon to Gold has been modeled using the Local Densitiy Approximation (LDA)

$$P_{LDA}(\mathbf{p}, E) = P_{MF}(\mathbf{p}, E) + P_{corr}(\mathbf{p}, E)$$



measured contribution corresponding to low momentum nucleons, occupying the shell model states

high momentum nucleons calculable using the result of uniform nuclear matter "recomputed" for a finite nucleus of mass number A

# The Relativistic Mean Field approximation

## already introduced in the M.B. Barbaro's talk model based on

M. Udias et al., Phys. Rev. C 64, 024614 (2001); C. Maieron et al., Phys. Rev. C 68, 048501 (2003); M. C. Martinezet al., Phys. Rev. C 73, 024607 (2006)

Still using the impulse approximation

■ The nuclear current is obtained as a sum over individual single-nucleon currents

$$J_N^{\mu}(\nu, \vec{q}) = \int d\vec{p} \, \bar{\psi}_F(\vec{p} + \vec{q}) \hat{J}_N^{\mu}(\nu, \vec{q}) \psi_B(\vec{p})$$

 $\psi_B$  = wave function for initial bound nucleons

 $\psi_F$  = wave function for final bound nucleons

$$\hat{J}^{\mu}_N(\nu,\vec{q}) = \text{relativistic nucleon current operator} = F_1(Q^2)\gamma_{\mu} + i\tfrac{k}{2m}F_2(Q^2)\sigma_{\mu\nu}q^{\nu} + \dots$$

 Matrix elements can be computed having the wave functions of the initial and the final nucleons (besides form factors)



# The Relativistic Mean Field approximation

- both bound and scattered nucleons feel the same 'potentials' which represent the nuclear medium;
- these potential are computed from lagrangians describing interactions among nucleons via boson exchange  $(\sigma, \omega)$ ;
- being a relativistic model,  $\psi_B$  and  $\psi_F$  are solutions of Dirac-like equations

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solving Dirac-like equations with scalar-vector (S-V) potentials:

$$\begin{split} \tilde{E}\gamma_0 - \vec{p} \cdot \vec{\gamma} - \tilde{M} &= 0 \\ \tilde{E} &= E - V(r) \\ \tilde{M} &= M - S(r) \end{split}$$

## The Relativistic Fermi Gas Model

- many MonteCarlo codes (GENIE, NuWro, Neut, Nuance) use some version of the Fermi model
  - target nucleons are moving (Fermi motion) subject to a nuclear potential (binding energy)
  - the ejected nucleon does not interact with other nucleons (Plane Wave Impulse Approximation)
  - Pauli blocking reduces the available phase space for scattered particle
- in terms of Spectral Function:

$$P_{RFGM} = \left(\frac{6\pi^2 A}{p_F^3}\right)\theta(p_F - \vec{p})\delta(E_{\vec{p}} - E_B + E)$$

### where

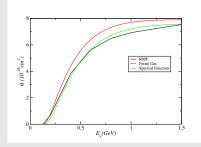
 $p_F = \text{Fermi momentum}$  (225 MeV for Oxygen)  $E_B = \text{average binding energy}$  (25 MeV for Oxygen)

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The  $\nu$ -nucleus cross sections

# The $\nu$ -nucleus cross sections ( $\nu A \rightarrow \mu X$ )

- **some** of the *qualitative* impacts of several nuclear models on the  $\nu$  observables can already be understood at the "cross section" level
- however the quantitative differences should be carefully evaluated



- as expected, FG overstimates the xsection over the whole QE energy regime
- $ightharpoonup m_A \sim 1 \text{ GeV}$  in any of the models
- dipole form factors
- **same** pattern for  $\bar{\nu}$

- Concept introduced by Zucchelli, Phys.Lett.B532:166-172,2002
- it involves producing a beam of  $\beta$ -unstable heavy ions (i.e.,  $^6$ He and  $^{18}$ Ne), accelerating them to some reference energy, and allowing them to decay in the straight section of a storage ring, resulting in a very intense  $\nu_e$  neutrino beam
  - lue pure u fluxes (e.g., only one neutrino species, in contrast to a conventional super-beam where contamination of other neutrino species is inevitable)
  - systematics free, since the spectrum can be calculated exactly (again, in contrast with a conventional beam, where knowledge of the spectrum always involves a sizable systematic uncertainty).

in the ion rest-frame:

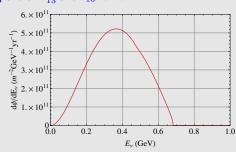
$$\frac{dN^{\rm rest}}{d\cos\theta dE_{\nu}} \sim E_{\nu}^2 (E_0 - E_{\nu}) \sqrt{(E_{\nu} - E_0)^2 - m_e^2}$$

in the laboratory frame:

$$\frac{d\Phi^{\text{lab}}}{dSdy}\bigg|_{\theta \simeq 0} \simeq \frac{N_{\beta}}{\pi L^2} \frac{\gamma^2}{g(y_e)} y^2 (1-y) \sqrt{(1-y)^2 - y_e^2}$$

# The βBeam concept

- lacktriangledown the value of the Lorentz boost factor  $\gamma$  and the source-detector distance Ldetermine the neutrino spectra
- interested in  $\nu_e \rightarrow \nu_\mu$  oscillation
- $\blacksquare$  leading terms in  $P_{\nu_e\nu_\mu}$  depend on  $\theta_{13}^2$  and  $\theta_{13}\cdot\sin\delta$
- here we focus on  $(\gamma, L) = (100, 732 \, Km)$
- $-(\nu \bar{\nu})$  spectra very similar
- QF events
- very low backgrounds



warning: working in the region where IA starts to be inadequate use this  $\beta$ Beam as a prototype!

# The CP discovery potential

## Definition

for any  $\theta_{13}$  is the ensemble of true values of  $\delta_{CP}$  for which the  $3\sigma$  CL do not touch  $\delta_{CP}=0,\pi,\pm\pi$ 

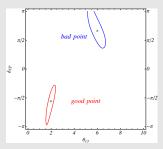
the precision measurement should be enough to establish  $\delta_{CP} \neq 0, \pi, \pm \pi$ 

# The CP discovery potential

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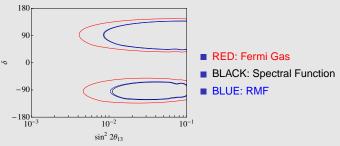
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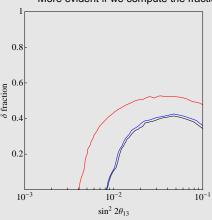
# The CP discovery potential

- We simulate at the same time  $\nu_e \rightarrow \nu_\mu$  and the CP-conjugate channel and compute event rates ( $\mu$  in the final state) after interaction with Oxygen
- "Points" inside the curves represent values of  $\delta_{CP}$  for which leptonic CP violation can be established at  $3\sigma$  CL



- 1- the FG performs too well compared with the other two models
- 2- at  $\delta \sim \pm \pi/2$  the largest discrepancy: 25-30% better!

More evident if we compute the fraction of  $good \delta$ 's over the total



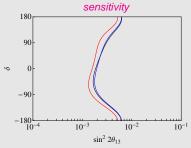
- RED: Fermi Gas
- BLACK: Spectral Function
- BLUE: RMF

# The sensitivity to $\theta_{13}$

 $\blacksquare$  same analysis for  $\theta_{13}$ 

## Definition

for any  $\delta_{CP}$  is the ensemble of true values of  $\theta_{13}$  for which the  $3\sigma$  CL do not touch  $\theta_{13} = 0$ 



- RFD: Fermi Gas
- BLACK: Spectral Function
- BLUE: RMF

A combined analysis

# A combined analysis

## What about a simultaneous fit to $\theta_{13}$ and $\delta_{CP}$ ?

To see the impact of various models:

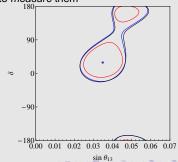
- we first fix some *true* value  $(\theta_{13}, \delta_{CP}) = (3^o, 30^o)$
- then we study the capability of the facility to measure them

RED: Fermi Gas

**BLACK: Spectral Function** 

**BLUE: RMF** 

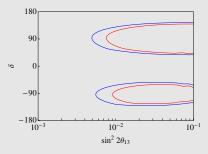
- much better precision at 3σCL for FG



Generalizing the previous results

# Generalizing the previous results

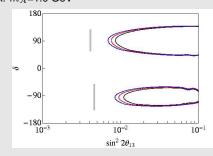
■ same effects with <sup>56</sup>Fe target blue: FG, red: SF



mild dependence on the axial mass

Facility and observables

SF with blue:  $m_A$ =1.2 GeV, red:  $m_A$ =1.1 GeV black:  $m_A$ =1.0 GeV



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# Facility a OO OOO O O

# Summary and conclusions

## Summary

- We studied the impact of nuclear effects on the determination of various neutrino parameters
- In particular, we compare the FG results (widely adopted in MonteCarlo codes) with the SF and RMF approaches
- The different behaviour of the cross sections translates into overstimated sensitivity to  $\theta_{13}$  and  $\delta_{CP}$
- Although we focused on Oxygen, the same pattern is observed for other nuclear targets

## Conclusions

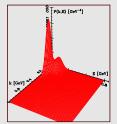
- It could be necessary to implement more realistic nuclear effects in MC codes
- It is also necessary to study the DIS region, where the future Neutrino Factories will work



### Benhar et al., Nucl. Phys. A 579 (1994) 493 Phys. Rev D72 (2005) 053005

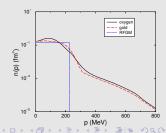
- overwhelming evidence from electron scattering that the energy-momentum distribution of nucleons in the nucleus is quite different from that predicted by Fermi gas
- the most important feature is the presence of strong nucleon-nucleon (NN) correlations (virtual scattering processes leading to the excitation of the participating nucleons to states of energy larger than the Fermi energy)

spectral function extends to  $|\mathbf{p}|\gg p_F$  and  $E\gg \varepsilon$ 



momentum distribution

$$n(\mathbf{p}) = \int dE \ P(\mathbf{p}, E)$$



$$\frac{d^2 \sigma_{\rm elem}}{d\Omega dE_l} = \frac{G_F^2 \, V_{ud}^2}{32 \, \pi^2} \, \frac{|k^{'}|}{|k|} \, \frac{1}{4 \, E_{\rm p} \, E_{|{\bf p}+{\bf q}|}} \, L_{\mu\nu} W^{\mu\nu}$$

■ The hadronic tensor is decomposed in structure functions as usual

$$\begin{array}{ll} W^{\mu\nu} & = & -g^{\mu\nu}\,W_1 + \tilde{p}^\mu\,\tilde{p}^\nu\,\frac{W_2}{m_N^2} + i\,\varepsilon_{\mu\nu\alpha\beta}\,\tilde{q}^\alpha\,\tilde{p}^\beta\,\frac{W_3}{m_N^2} + \tilde{q}^\mu\,\tilde{q}^\nu\,\frac{W_4}{m_N^2} + \\ & & (\tilde{p}^\mu\,\tilde{q}^\nu + \tilde{p}^\nu\,\tilde{q}^\mu)\,\frac{W_5}{m_N^2} \end{array}$$

 $\blacksquare$  the formalism can be applied to  ${\bf both}$  elastic and anelastic processes specifying the form of the structure functions  $W_i$ 

