

The Impact of nuclear effects on accelerator-based neutrino-oscillation physics

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[Repetition helps]

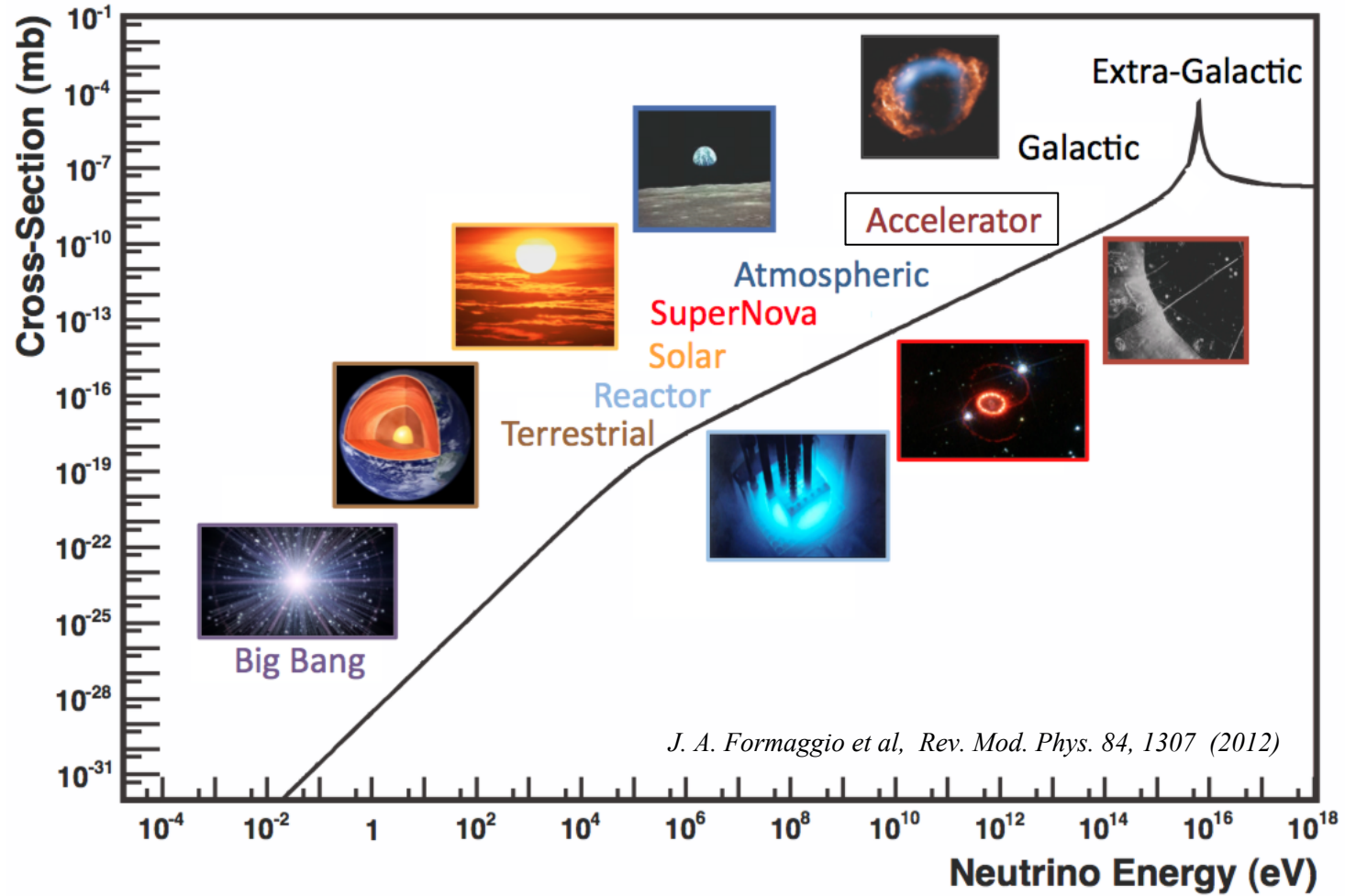


[Repetition helps]

SED STUFANT

[But it's boring]

Neutrinos Everywhere!



Accelerator-based neutrino-oscillation experiments

Current knowledge:

	θ_{12}	θ_{13}	θ_{23}	$\Delta m_{21}^2/10^{-5}$	$\Delta m_{3j}^2/10^{-3}$	δ_{CP}
Normal Ordering	$33.56^{+0.77}_{-0.75}$	$8.46^{+0.15}_{-0.15}$	$41.6^{+1.5}_{-1.2}$	$7.50^{+0.19}_{-0.17}$	$2.524^{+0.039}_{-0.040}$	261^{+51}_{-59}
Inverted Ordering	$33.56^{+0.77}_{-0.75}$	$8.49^{+0.15}_{-0.15}$	$50.0^{+1.1}_{-1.4}$	$7.50^{+0.19}_{-0.17}$	$-2.514^{+0.038}_{-0.041}$	277^{+40}_{-46}

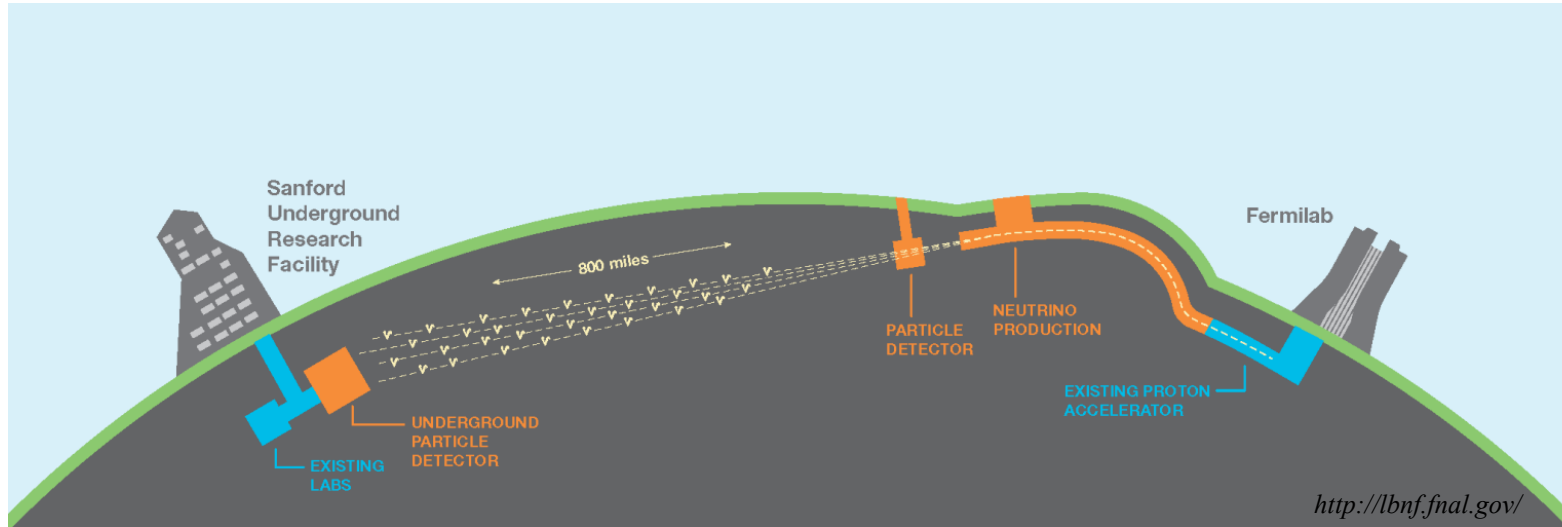
Current and future goals:

- Establish whether there is CP violation in the leptonic sector and, if so, measure δ_{CP}
- Improve the accuracy on θ_{23}
- Determine the neutrino mass ordering: $m_1 < m_2 < m_3$ or $m_3 < m_1 < m_2$

Many experiments:

- **MiniBooNE** (concluded), **NOvA** (running), etc.
- **SBN Program: MicroBooNE** (running), **ICARUS** (under construction), **SBND** (under construction); **DUNE** (under construction)
- **T2K** (running), **T2HK** (under construction)

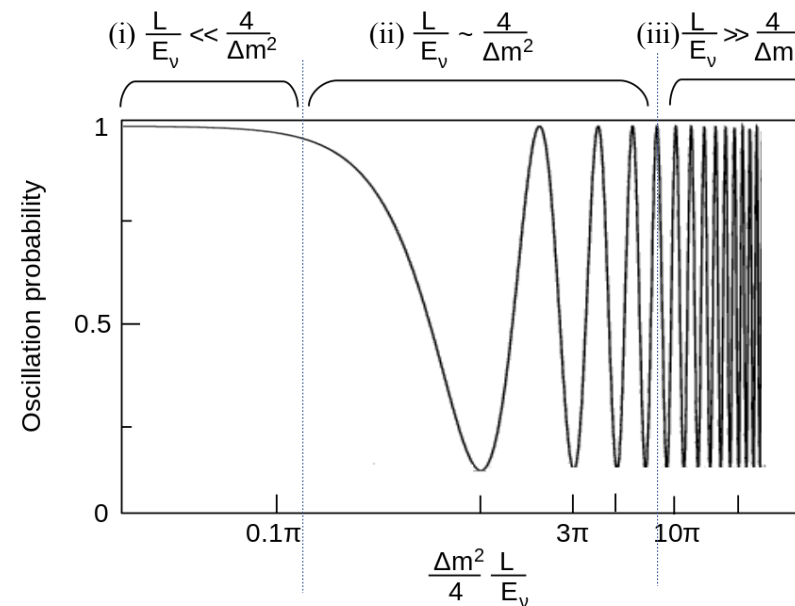
Accelerator-based neutrino-oscillation experiments



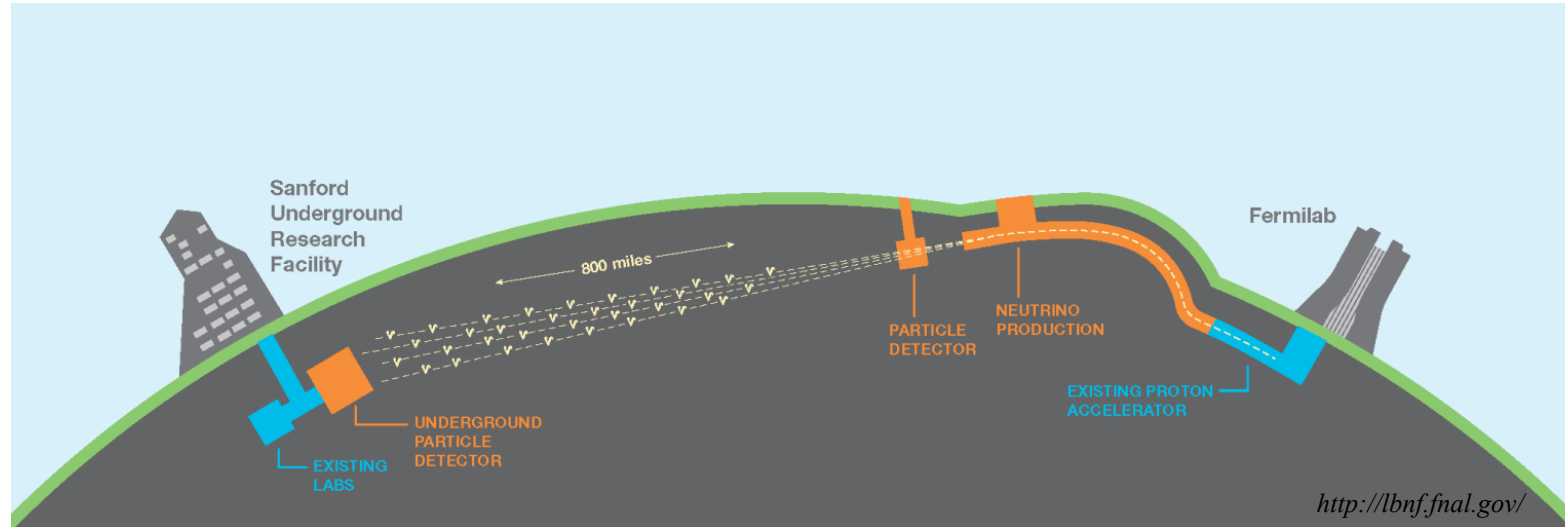
Oscillation Probability:

$$P(\nu_i \rightarrow \nu_j) = \sin^2 2\theta \sin^2 \left(\frac{\Delta m^2 L}{4E_\nu} \right)$$

- (i) Detector is too close to the source, no oscillation happened.
- (ii) Necessary but most sensitive region to detect oscillations.
- (iii) Several oscillations happened, only average oscillation probability can be measured.



Accelerator-based neutrino-oscillation experiments



Oscillation Probability:

$$P(\nu_i \rightarrow \nu_j) = \sin^2 2\theta \sin^2 \left(\frac{\Delta m^2 L}{4E_\nu} \right)$$

Maximal sensitivity to oscillation probability:

- Maximize the baseline L \longrightarrow Neutrino beam diverges
- Minimize the beam energy E_ν \longrightarrow Cross section decreases

Optimal combination:

$L \sim$ few 100s of km
 $E_\nu \sim$ Intermediate energies

Event Rates:

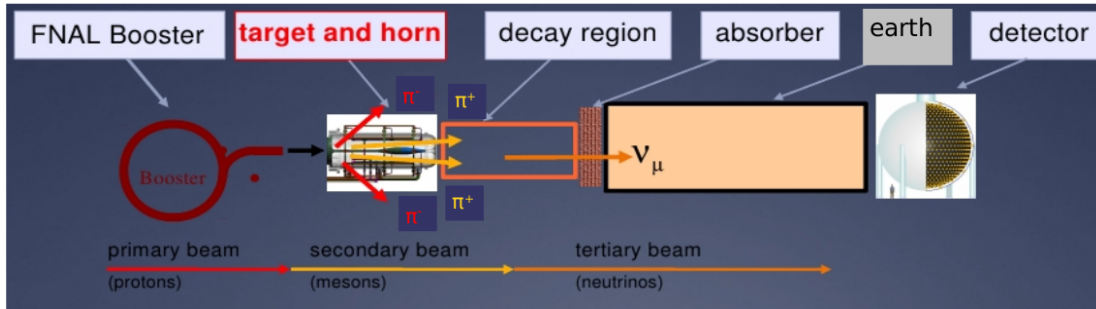
$$N_{\text{FD}}^{\alpha \rightarrow \beta}(\mathbf{p}_{\text{reco}}) = \sum_i \phi_\alpha(E_{\text{true}}) \times P_{\alpha\beta}(E_{\text{true}}) \times \sigma_\beta^i(\mathbf{p}_{\text{true}}) \times \epsilon_\beta(\mathbf{p}_{\text{true}}) \times R_i(\mathbf{p}_{\text{true}}; \mathbf{p}_{\text{reco}})$$

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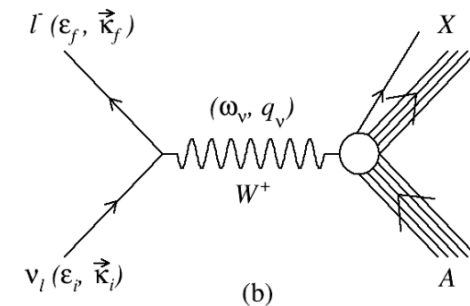
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Energy reconstruction:



- Neutrino beams are part of tertiary beam – neutrino energy is not know.
- The neutrino energy is reconstructed based on the kinematics of the final state particles.
- Using the **nuclear model implemented in the Monte Carlo generator.**

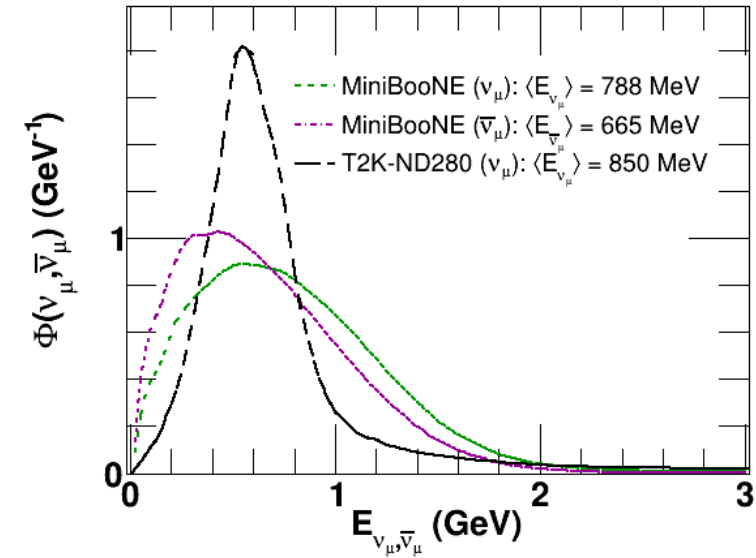
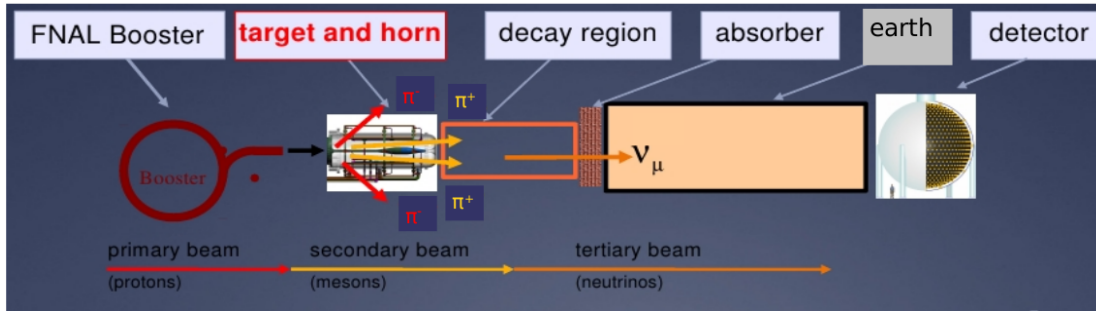


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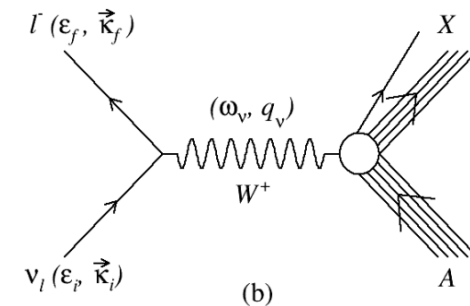
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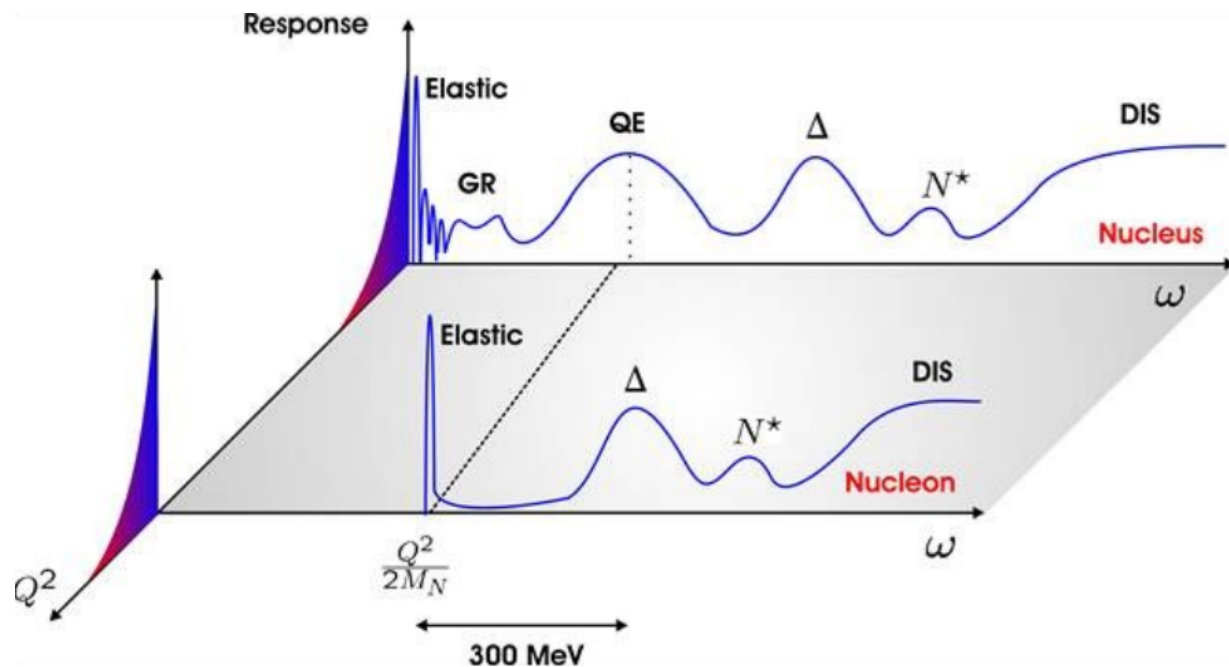


Cross section:

Event Rates:

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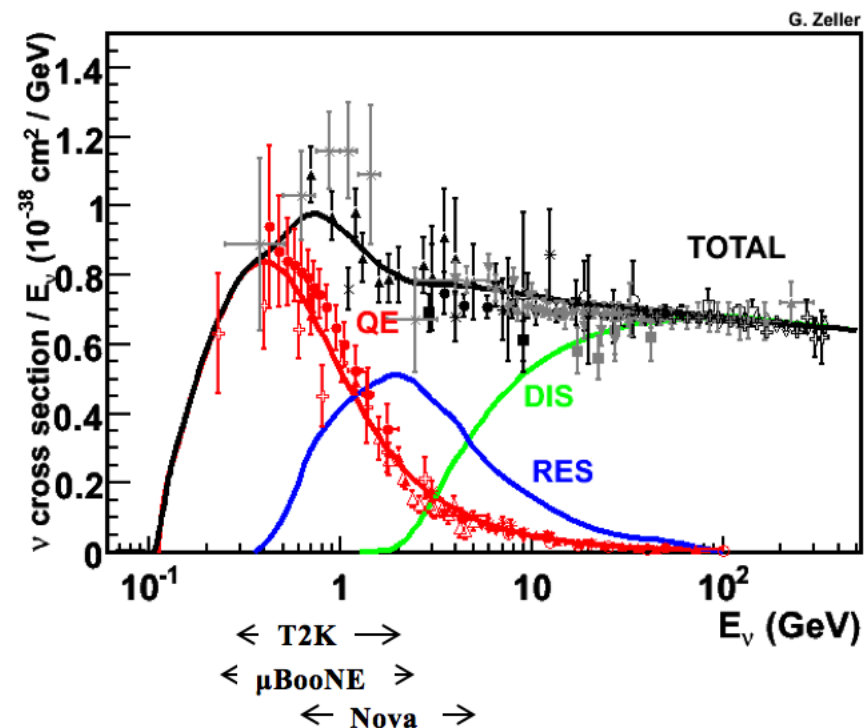
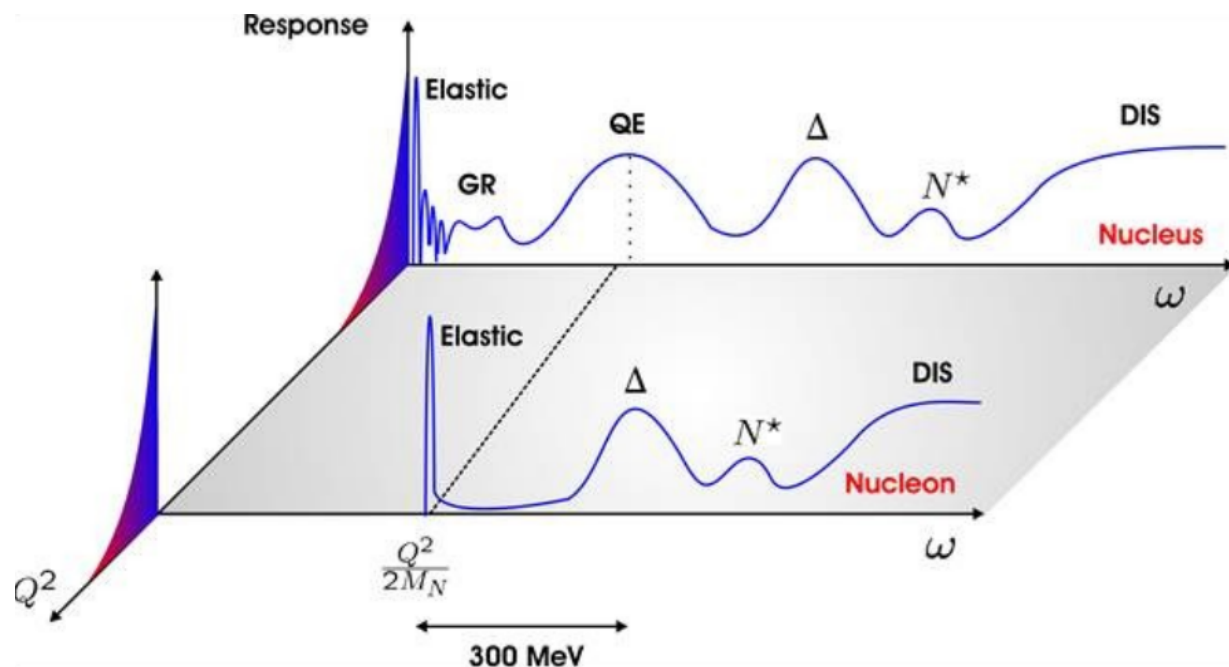
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Event Rates:

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Cross section:



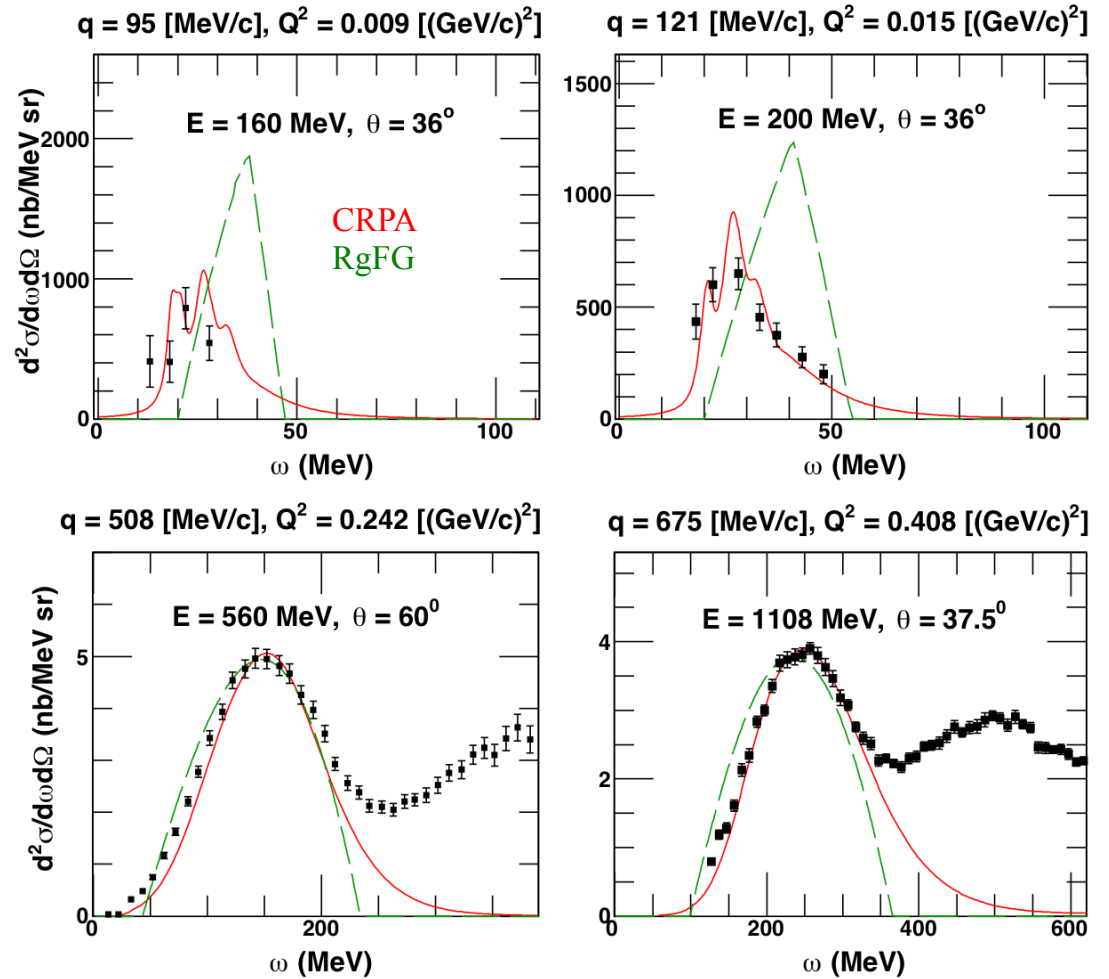
- Cross section is calculated using the **nuclear model implemented in the Monte Carlo generator.**

- Monte Carlo generators should consist of a realistic nuclear model that can describe accurately the neutrino-nucleus scatterings at the kinematics relevant for these experiments.
- Current status: Monte Carlo generators, e.g. GENIE, are primarily based on Fermi-gas based Model of the nucleus.

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Example from $^{12}\text{C} (e, e')$

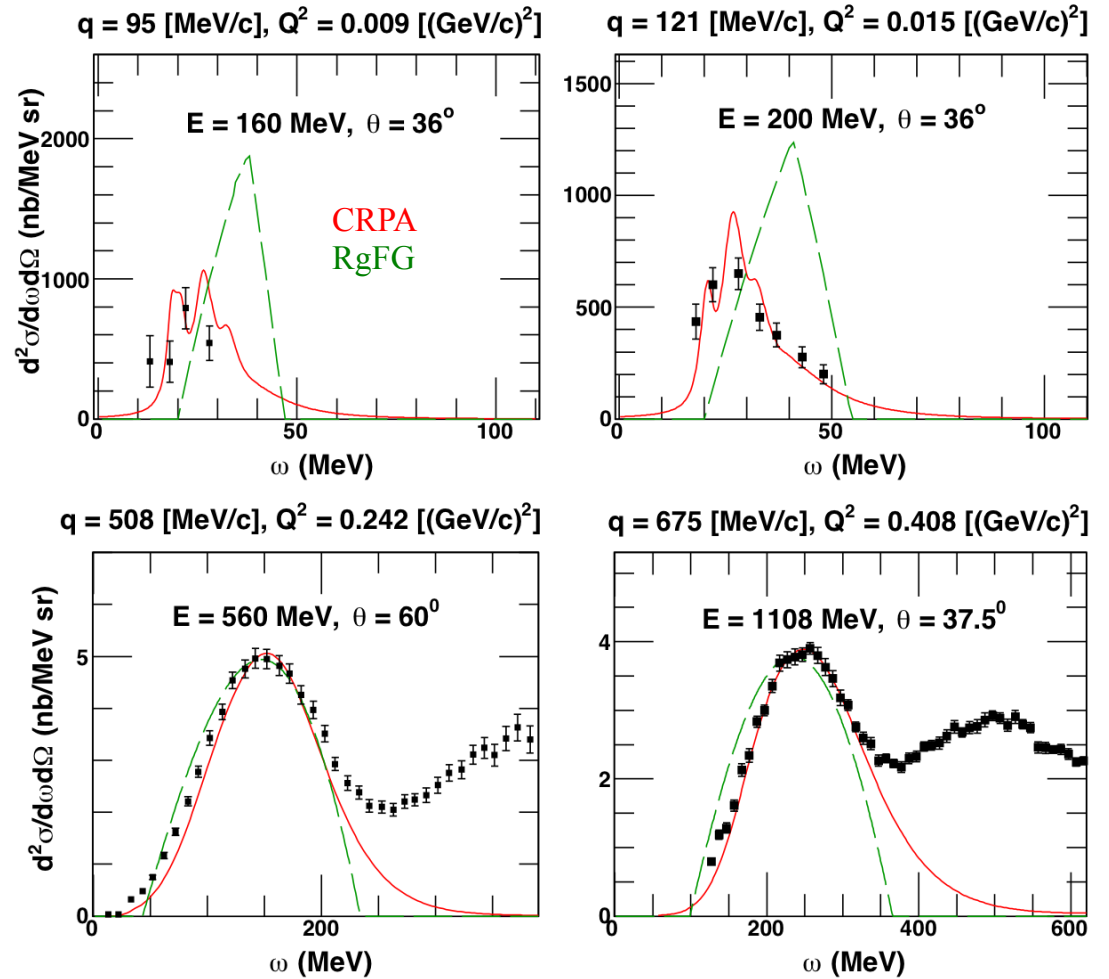


VP, PhD thesis (2016)

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Example from $^{12}\text{C} (e, e')$



- The uncertainties in the measurements, mainly owing to inadequate nuclear models in generators, is $\sim 10\text{-}15\%$ - No Surprise!**

Neutrino Experimentalist's Desk

ARE YOU COMING TO BED?

I CAN'T. THIS IS IMPORTANT.

WHAT?

MEASURING →

NEUTRINO CROSS SECTIONS

↓

- Need better models to reduce systematic uncertainties to eventually make precise measurements

Nuclear Theorist's Desk

ARE YOU COMING TO BED?

I CAN'T. THIS IS IMPORTANT.

WHAT?

← CALCULATING

NEUTRINO CROSS SECTIONS

↓

- Need precise data to test and validate the assumptions and approximations to eventually improve the model



- **In an ideal world:** with strong collaboration between neutrino experimentalists and nuclear theorists - we all understand neutrino cross section with high accuracy and make precise oscillations measurements.

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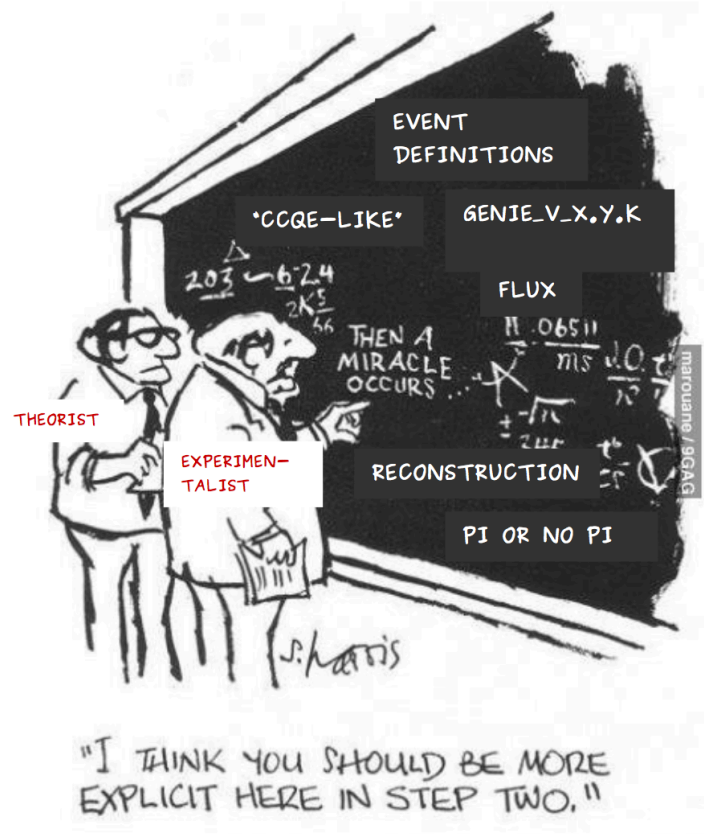
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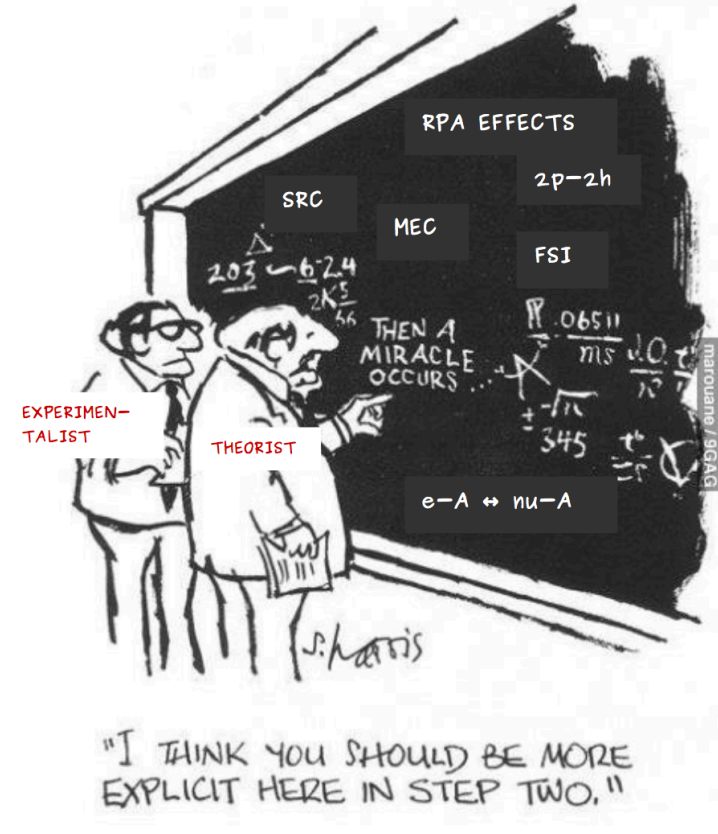
And they lived happily ever after...

- In real world: things are a bit messy!

When nuclear theorists talk to neutrino experimentalists



When neutrino experimentalists talk to nuclear theorists



State of art, unfortunately :-/

My Neutrino Experimentalist's Desk



My Nuclear Theorist's Desk



There's no such thing as multi-tasking – just doing lots of things badly. The correct term is multi-failing.

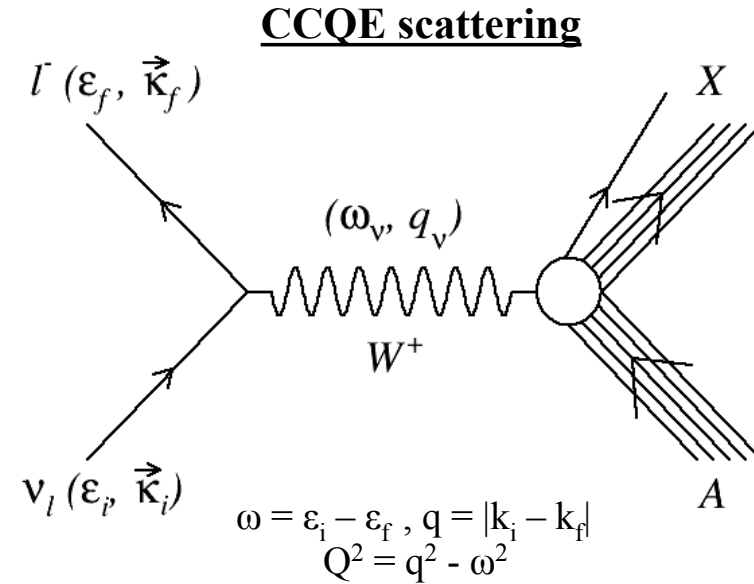
General cross section formalism:

Leptonic current

$$\begin{aligned} J_\mu &= \bar{\nu}_l \gamma_\mu (1 - \gamma_5) l \\ &= \bar{\nu}_l \gamma_\mu l - \bar{\nu}_l \gamma_\mu \gamma_5 l \end{aligned}$$

Hadronic current

$$\begin{aligned} J^\mu &= \bar{u}_N \left[\gamma^\mu F_1(Q^2) + \frac{i}{2M_N} \sigma^{\mu\nu} q_\nu F_2(Q^2) \right. \\ &\quad \left. + \gamma^\mu \gamma_5 F_A(Q^2) + \frac{1}{2M_N} q^\mu \gamma_5 F_P(Q^2) \right] u_N \end{aligned}$$



Propagator

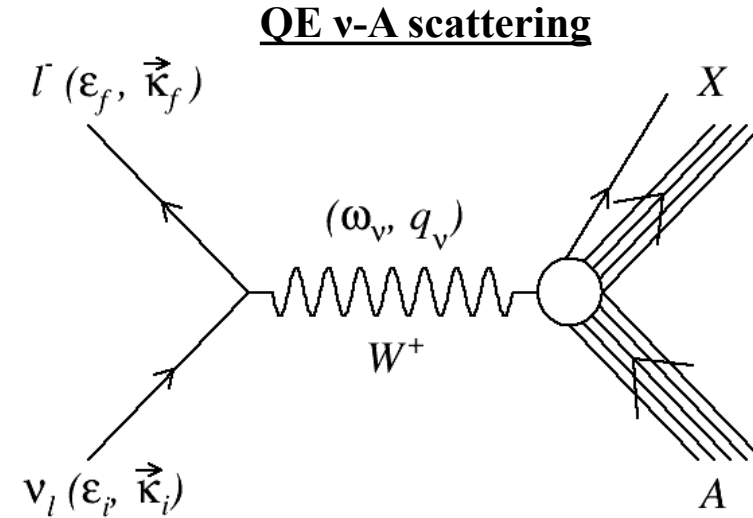
$$\approx \frac{1}{Q^2 - M_W^2}, \quad Q^2 \ll M_W^2$$

$$\left(\frac{d^2\sigma}{d\omega d\Omega} \right)_\nu = \frac{G_F^2 \cos^2 \theta_c}{(4\pi)^2} \left(\frac{2}{2J_i + 1} \right) \varepsilon_f \kappa_f \zeta^2(Z', \varepsilon_f, |q|) \left[\sum_{J=0}^{\infty} \sigma_{CL,\nu}^J + \sum_{J=1}^{\infty} \sigma_{T,\nu}^J \right]$$

$$\sigma_{CL,\nu}^J = [v_\nu^{\mathcal{M}} R_\nu^{\mathcal{M}} + v_\nu^{\mathcal{L}} R_\nu^{\mathcal{L}} + 2 v_\nu^{\mathcal{ML}} R_\nu^{\mathcal{ML}}] \quad \sigma_{T,\nu}^J = [v_\nu^T R_\nu^T \pm 2 v_\nu^{TT} R_\nu^{TT}]$$

+/- sign is the only difference between ν and anti- ν

General cross section formalism:



$$\left(\frac{d^2\sigma}{d\omega_\nu d\Omega} \right)_\nu = \frac{G_F^2 \cos^2 \theta_c}{(4\pi)^2} \left(\frac{2}{2J_i + 1} \right) \varepsilon_f \kappa_f$$

$$\times \zeta^2(Z', \varepsilon_f, q_\nu) \left[\sum_{J=0}^{\infty} \sigma_{CL,\nu}^J + \sum_{J=1}^{\infty} \sigma_{T,\nu}^J \right]$$

$$\sigma_{CL,\nu} = [v_\nu^M R_\nu^M + v_\nu^L R_\nu^L + 2 v_\nu^{ML} R_\nu^{ML}]$$

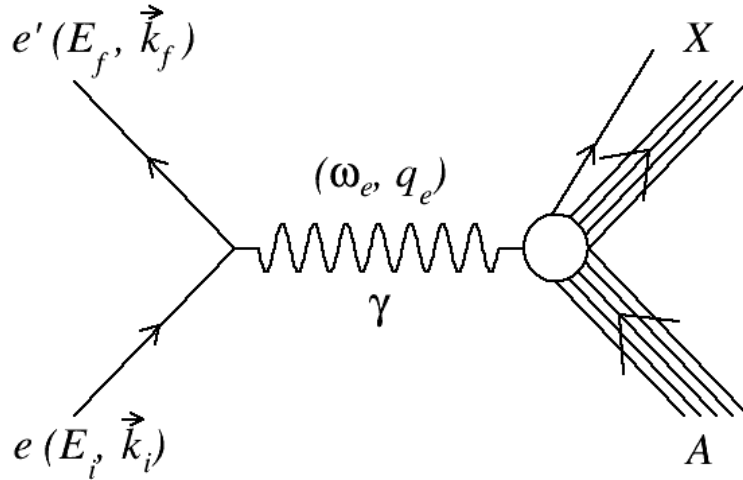
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v 's \rightarrow Leptonic coefficients \rightarrow Purely kinematical \rightarrow Easy to calculate

R 's \rightarrow Response functions \rightarrow Nuclear dynamics \rightarrow **Need nuclear models to calculate!**

General cross section formalism:

QE e-A scattering

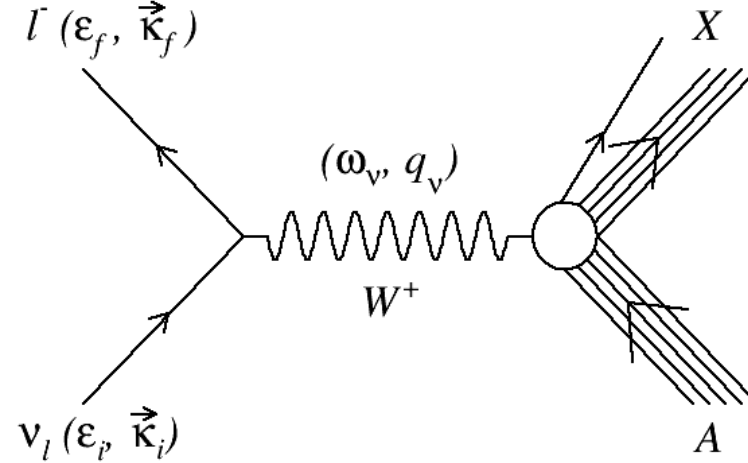


$$\left(\frac{d^2\sigma}{d\omega_e d\Omega} \right)_e = \frac{\alpha^2}{Q^4} \left(\frac{2}{2J_i + 1} \right) \frac{1}{k_f E_i} \times \zeta^2(Z', E_f, q_e) \left[\sum_{J=0}^{\infty} \sigma_{L,e}^J + \sum_{J=1}^{\infty} \sigma_{T,e}^J \right]$$

$$\sigma_{L,e} = v_e^L R_e^L$$

$$\sigma_{T,e} = v_e^T R_e^T$$

QE v-A scattering



$$\left(\frac{d^2\sigma}{d\omega_\nu d\Omega} \right)_\nu = \frac{G_F^2 \cos^2 \theta_c}{(4\pi)^2} \left(\frac{2}{2J_i + 1} \right) \varepsilon_f \kappa_f \times \zeta^2(Z', \varepsilon_f, q_\nu) \left[\sum_{J=0}^{\infty} \sigma_{CL,\nu}^J + \sum_{J=1}^{\infty} \sigma_{T,\nu}^J \right]$$

$$\sigma_{CL,\nu} = [v_\nu^M R_\nu^M + v_\nu^L R_\nu^L + 2 v_\nu^{ML} R_\nu^{ML}]$$

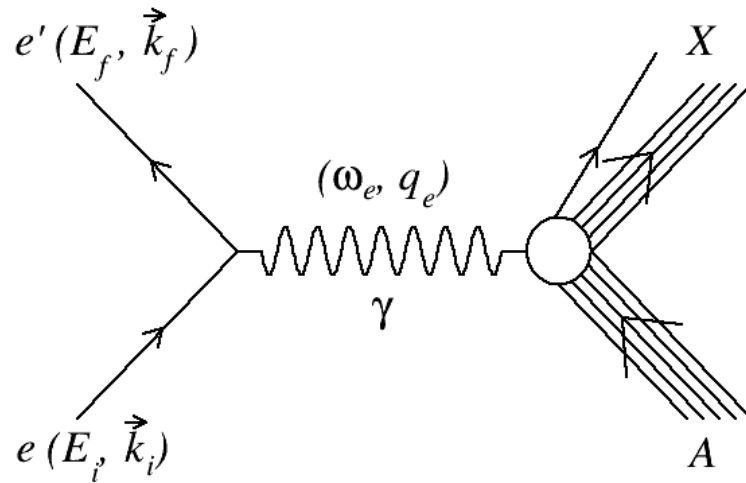
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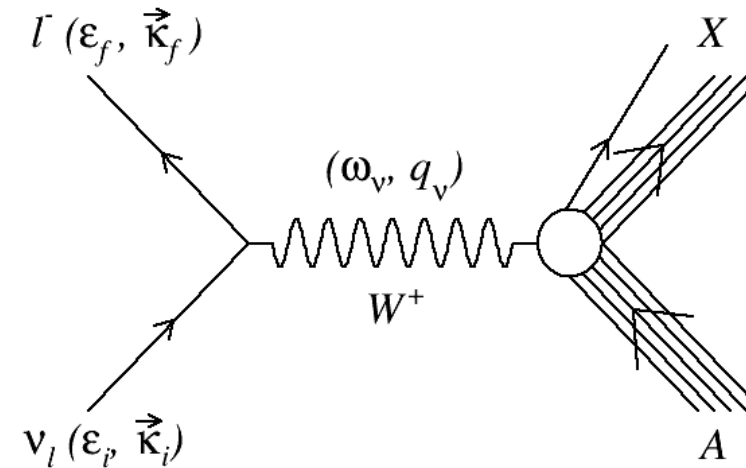
R 's \rightarrow Response functions \rightarrow Nuclear dynamics \rightarrow **Need nuclear models to calculate!**

e-A to v-A cross section

QE e-A scattering



QE v-A scattering



CAUTION: CHALLENGES AHEAD

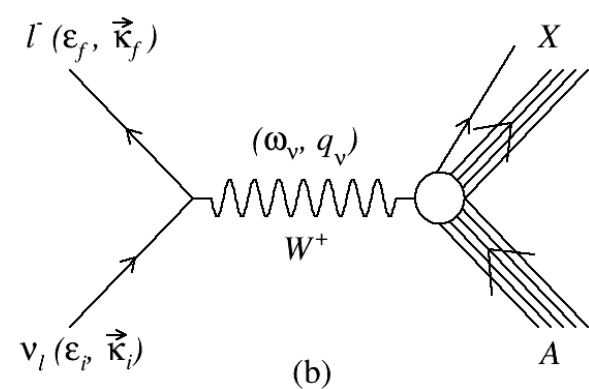
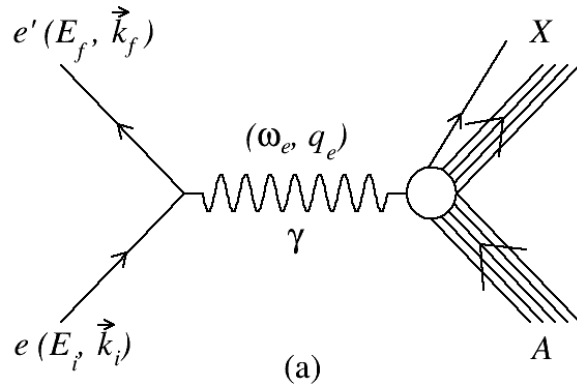
- Monoenergetic beams (E_e known precisely; ω, q can be calculated)
- Hence, different reaction channels can be separated

- E_ν has to be reconstructed \rightarrow wide flux
- Hence, difficult to distinguish different reactions channels.
- What experimentalist's measure:

$$\frac{d^2\sigma}{dT_l \cos\theta} = \frac{\sum_j U_{ij}(d_j - b_j)}{\Phi \cdot T \cdot \epsilon_i \cdot (\Delta T_l, \Delta \cos\theta)_i}$$

- What theorist's calculate:

$$\frac{d^2\sigma}{dT_l d \cos\theta} = \frac{1}{\int \Phi(E_\nu) dE_\nu} \int dE_\nu \left[\frac{d^2\sigma}{d\omega d \cos\theta} \right]_{\omega=E_\nu - E_l} \Phi(E_\nu)$$



> Start with some basic description of nucleus. Add some ingredient, make some approximation, etc. to take into account several nuclear effects.



> Test each of those ingredient, and approximation against e-A scattering data*.



> *Congratulations madam, you rock! your model is validated. Please extended the model (with few relatively simpler steps) into nu-A scattering.*



> If it all successfully describes the e-A scattering data over wide range of kinematics (and possibly over several nuclei).



> Implement the model in MC. Validate implementation against e-A data.



> **Ready to use by experiments!**



> If it doesn't work for e-A scattering (*sorry sir, you suck!*), its highly unlikely to work for nu-A scattering.

*There is plenty of data available for e-A scattering over wide range of kinematics and for several nuclei.

Few (many) models:

- RFG+RPA (Martini, Ericson, *et al*)
- RFG+RPA (Nieves, *et al*)
- HF+CRPA (Pandey, Jachowicz *et al*)
- Spectral Function Formalism (Benhar, Rocco, Ankowski, *et al.*)
- Super-Scaling approach (Amaro, Barbaro, Caballero, Donnelly, Megias, *et al.*)
- Relativistic Green's Function Model (Meucci, Giusti *et al.*)
- Green's Function Monte Carlo Approach (Lovato, Gandolfi, Carlson, *et al.*)
- ..., *etc.*

Formalism: HF-CRPA Approach

References:

PRC 89, 024601 (2014); PRC 92, 024606 (2015); PRC 94, 015501 (2016);
PRC 94, 024611 (2016); PRC 94, 054609 (2016);
arXiv:1612.05511 [nucl-th] (accepted in PRD); arXiv:1704.07817 [nucl-th]

- We start by describing the nucleus with a Hartree-Fock (HF) approximation. The mean-field (MF) potential is obtained by solving the HF equations and using a Skyrme (SkE2) two-body interaction.
- Once we have bound and continuum single-nucleon wave functions, we introduce long-range correlations between the nucleons through a continuum Random Phase Approximation (CRPA).
- The propagation of particle-hole pairs in the nuclear medium is described by the polarization propagator. In the Lehmann representation, this particle-hole Green's function is given by

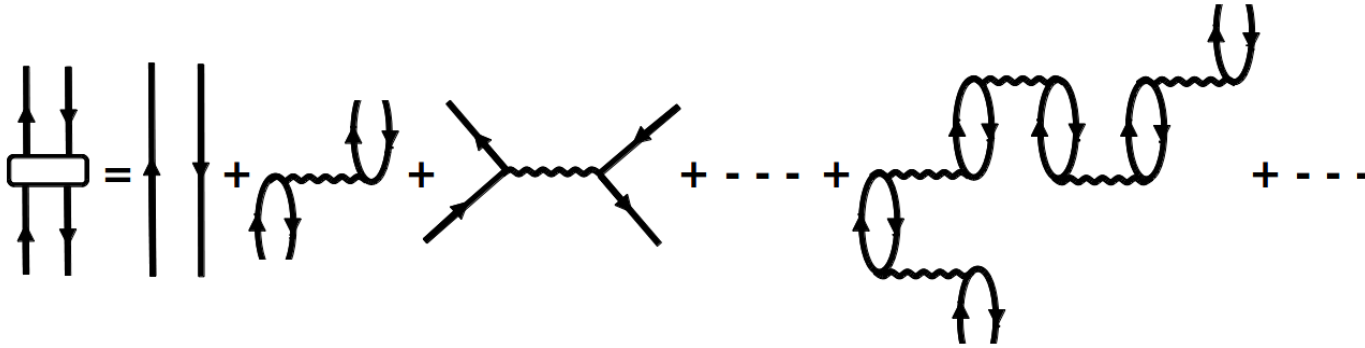
$$\Pi(x_1, x_2, x_3, x_4; E_x) = \hbar \sum_n \left[\frac{\langle \Psi_0 | \hat{\psi}^\dagger(x_2) \hat{\psi}(x_1) | \Psi_n \rangle \langle \Psi_n | \hat{\psi}^\dagger(x_3) \hat{\psi}(x_4) | \Psi_0 \rangle}{E_x - (E_n - E_o) + i\eta} - \frac{\langle \Psi_0 | \hat{\psi}^\dagger(x_3) \hat{\psi}(x_4) | \Psi_n \rangle \langle \Psi_n | \hat{\psi}^\dagger(x_2) \hat{\psi}(x_1) | \Psi_0 \rangle}{E_x + (E_n - E_o) - i\eta} \right]$$

- RPA equations are solved using a Green's function approach.

Formalism: HF-CRPA Approach

- The RPA-polarization propagator

$$\begin{aligned}\Pi^{(RPA)}(x_1, x_2; E_x) &= \Pi^{(0)}(x_1, x_2; E_x) + \frac{1}{\hbar} \int dx dx' \Pi^{(0)}(x_1, x; E_x) \\ &\quad \times \tilde{V}(x, x') \Pi^{(RPA)}(x', x_2; E_x)\end{aligned}$$



- The Skyrme (SkE2) nucleon-nucleon interaction, which was used in the HF calculations, is also used to perform CRPA calculations. That makes this approach self-consistent.

Formalism: Folding procedure

A limitation of RPA formalism at lower energies:

- energy position of the giant resonances is generally well predicted
- width is underestimated
- height is overestimated



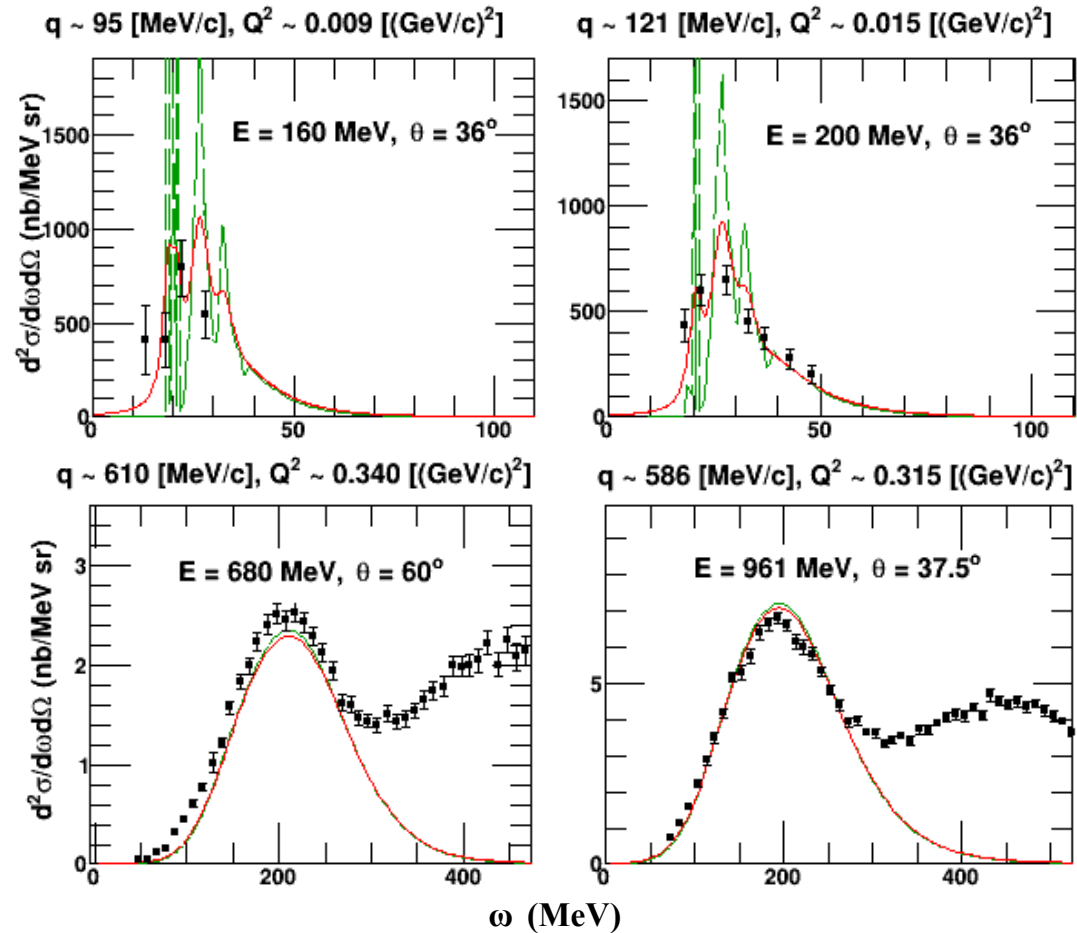
Folding

$$R'(q, \omega') = \int_{-\infty}^{\infty} d\omega R(q, \omega) L(\omega, \omega')$$

$$L(\omega, \omega') = \frac{1}{2\pi} \left[\frac{\Gamma}{(\omega - \omega')^2 + (\Gamma/2)^2} \right]$$

$$\Gamma = 3 \text{ MeV}$$

CRPA: Without Folding, CRPA: With Folding



Formalism: Relativistic corrections

Non-relativistic model



'Semi-relativistic' modeling

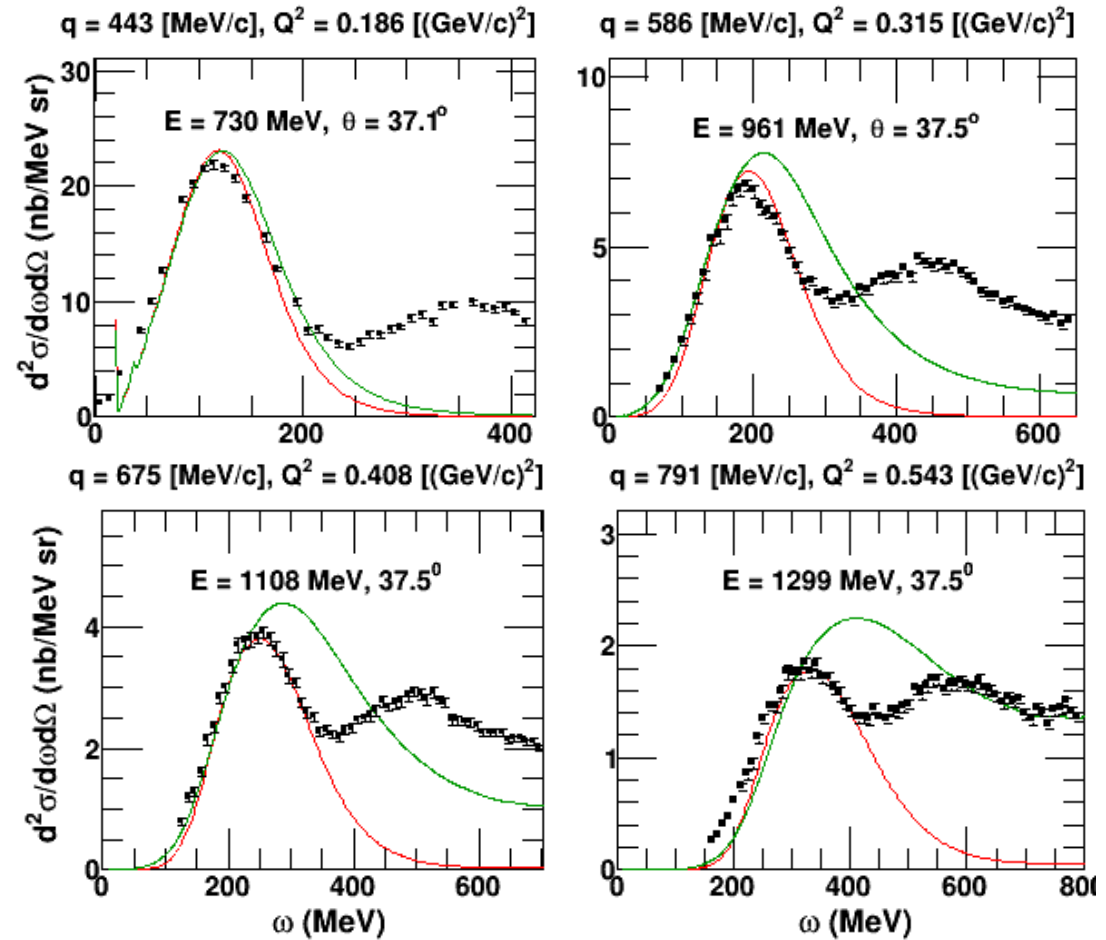
(S. Jeschonnek and T. W. Donnelly, PRC57, 2438 (1998))

Kinematic effects:

$$\lambda \rightarrow \lambda (1+\lambda), \quad \lambda = \omega/2M_N$$

$$\text{QE peak: } \omega = q^2/2M_N \Rightarrow Q^2/2M_N$$

CRPA: Non-relativized, CRPA: Relativized



Formalism: Regularization of the residual interaction

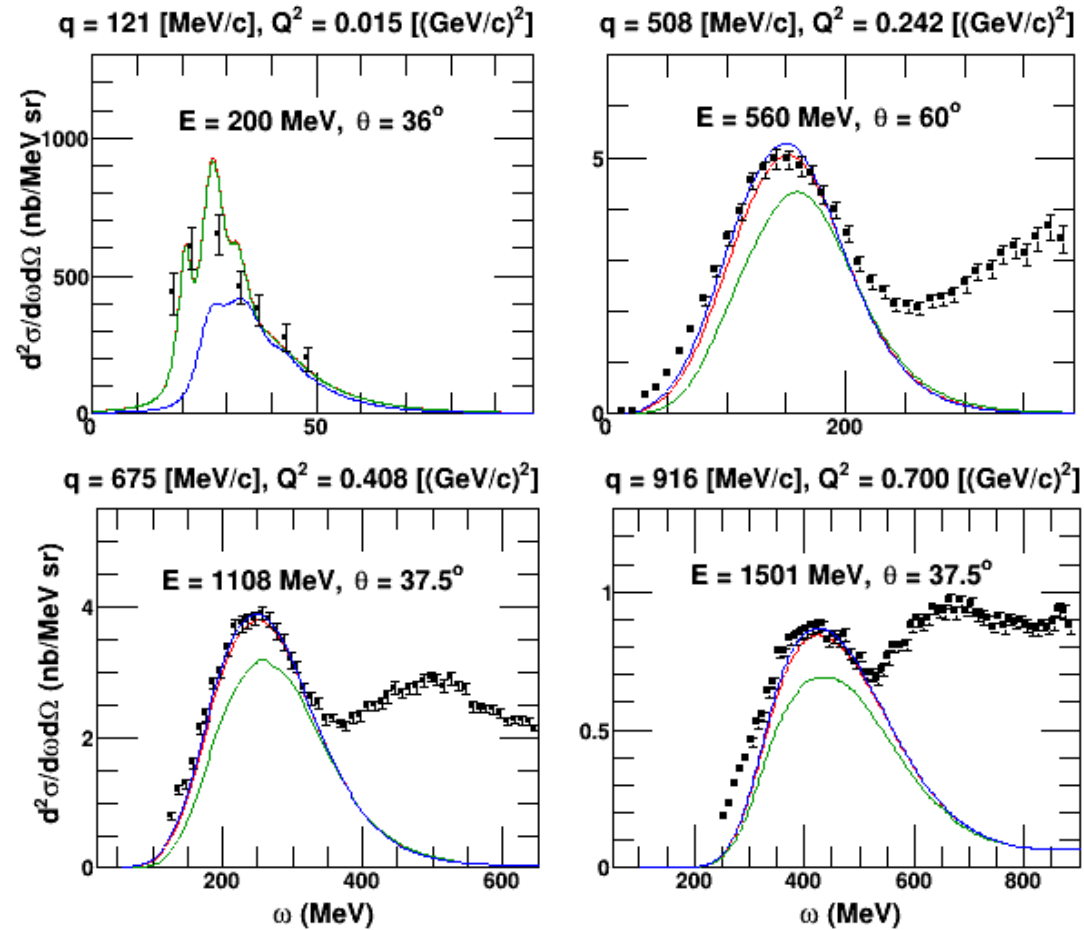
SkE2 interaction → optimized against ground-state and low-excitation energy properties
→ at higher Q^2 , unrealistically strong!



$$V(Q^2) = V(Q^2 = 0) \frac{1}{\left(1 + \frac{Q^2}{\Lambda^2}\right)^2}$$

With $\Lambda = 455$ MeV, optimized with a χ^2 fitting of theory-experiment comparison from low ω up to the QE peak, over broad set of available data on $A(e,e')$ scattering.

CRPA: without dipole, CRPA: with dipole, HF



Formalism: Coulomb correction

Effect of coulomb potential of the nucleus on the charged lepton:

→ Low energies: Fermi function $F(Z', E) = \frac{2\pi\eta}{1 - e^{-2\pi\eta}}$ $\eta \sim \mp Z' \alpha$

→ High energies:

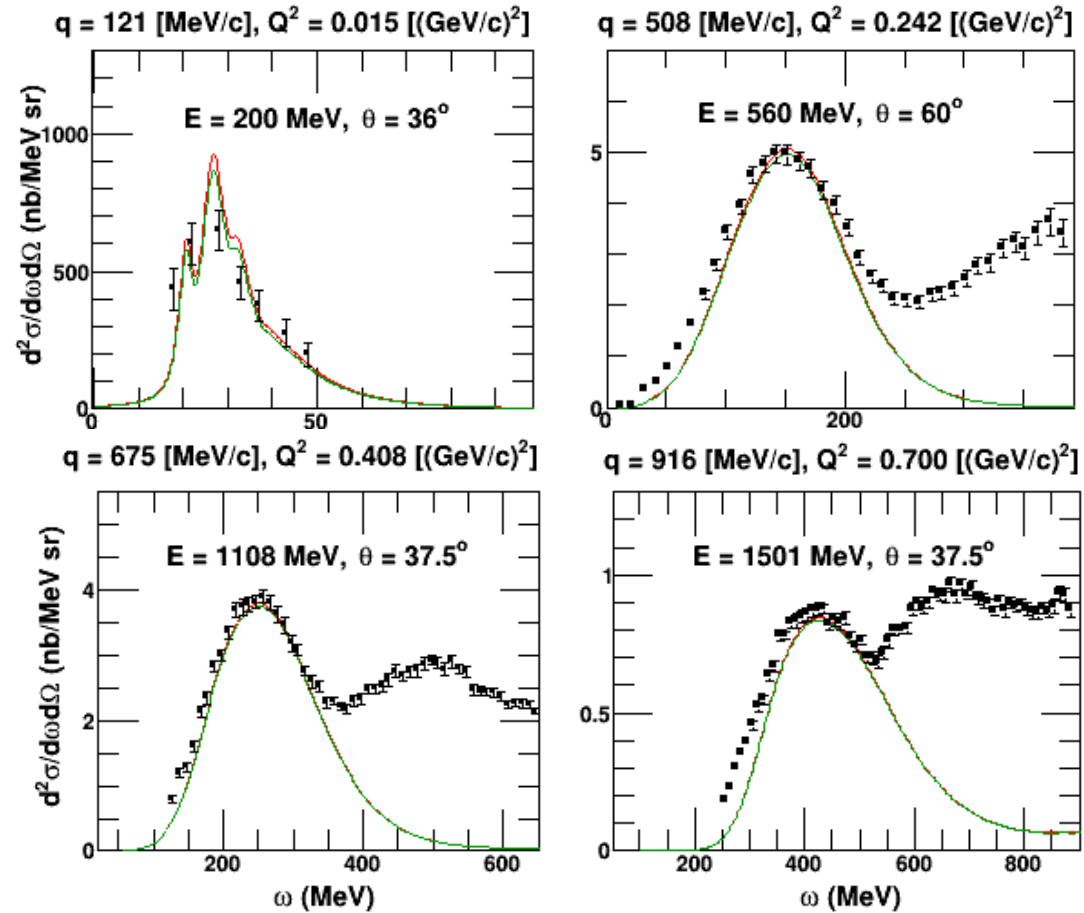
CRPA: without MEMA, CRPA: with MEMA

**Modified effective
momentum approximation**

(J. Engel, PRC57,2004 (1998))

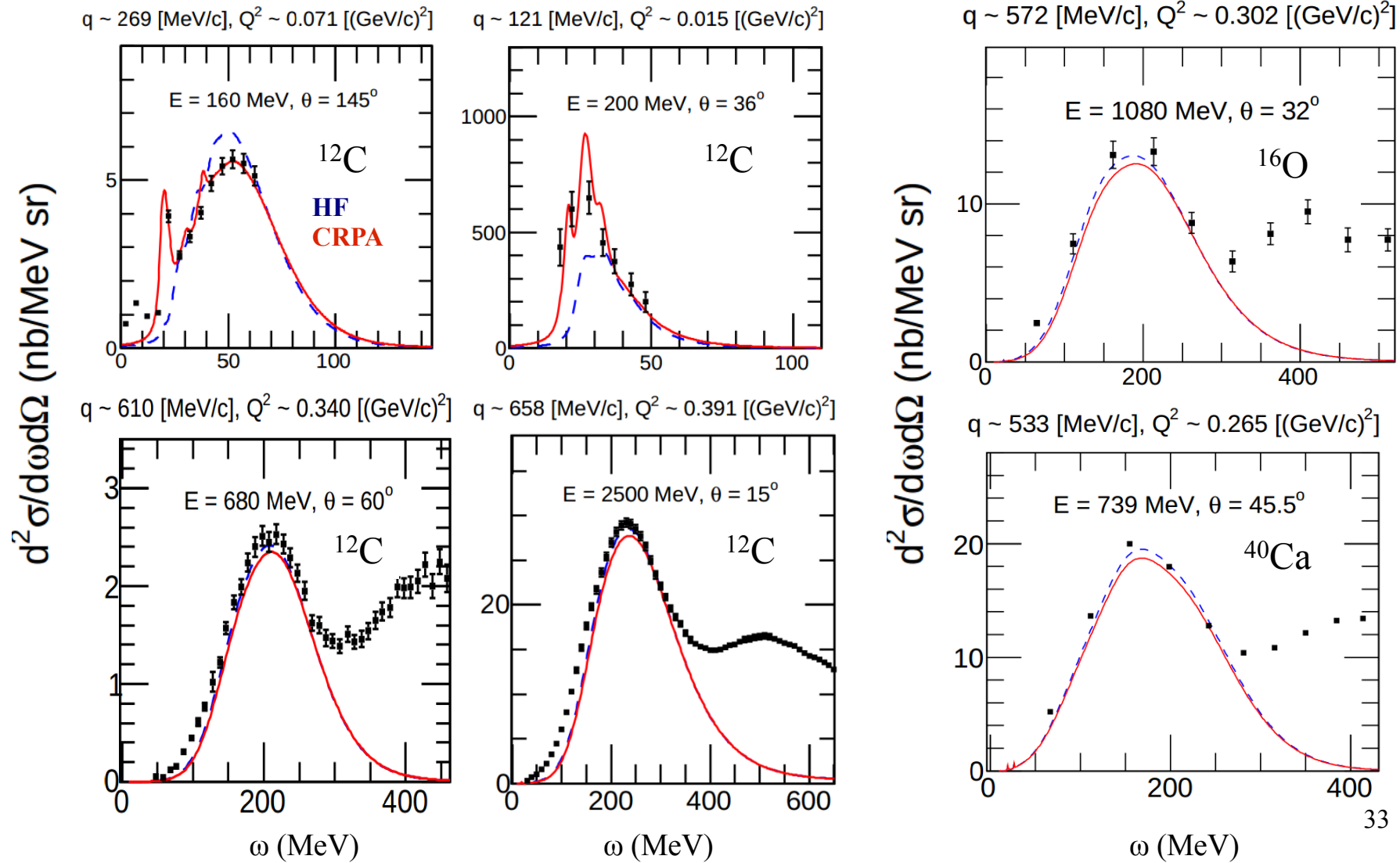
$$q_{\text{eff}} = q + 1.5 \left(\frac{Z' \alpha \hbar c}{R} \right)$$

$$\zeta(Z', E, q) = \sqrt{\frac{q_{\text{eff}} E_{\text{eff}}}{q E}}$$

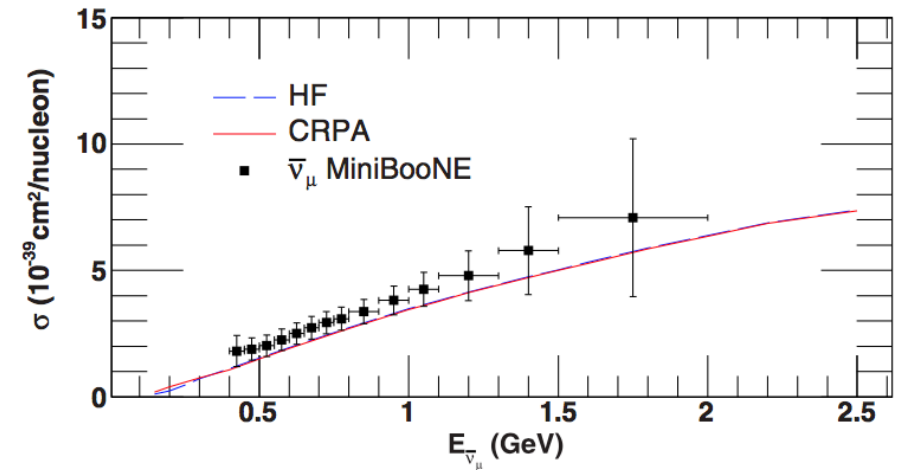
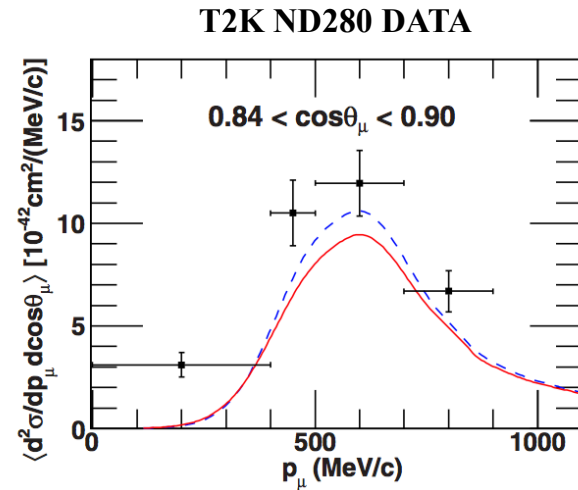
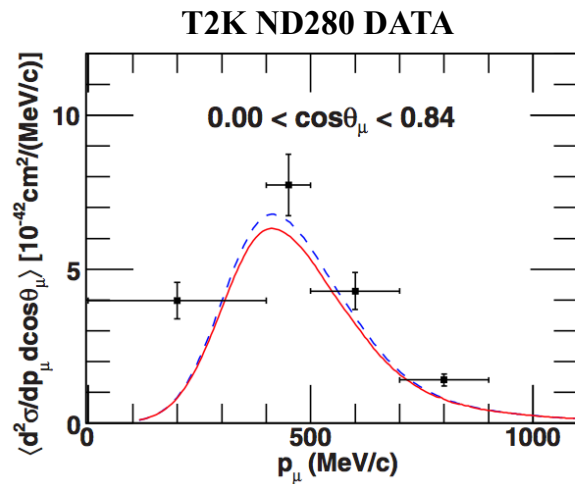
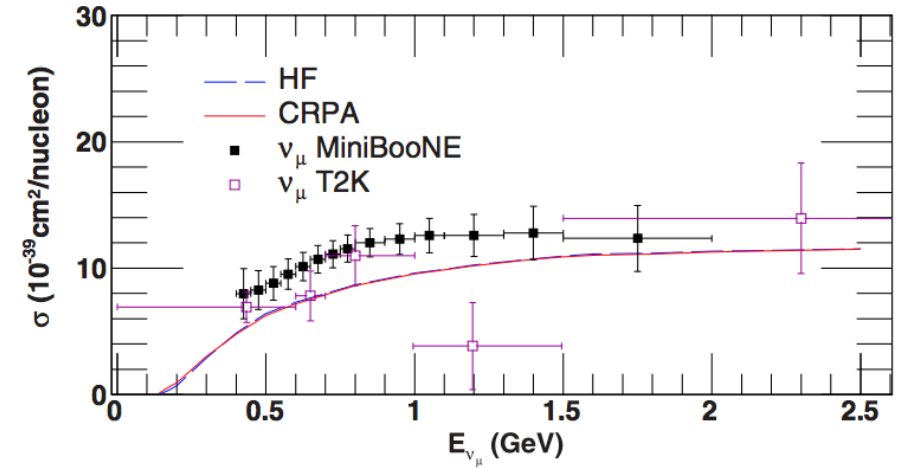
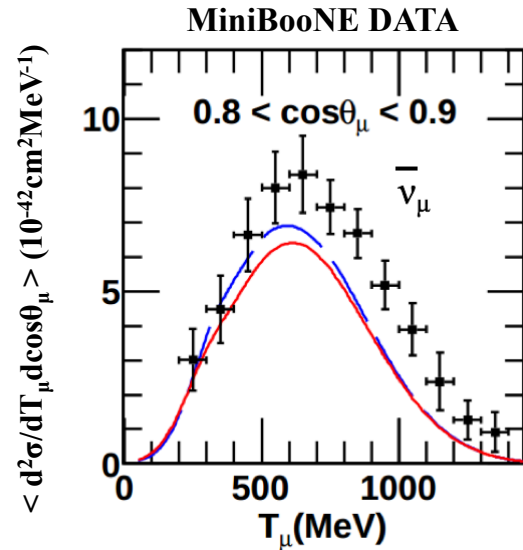
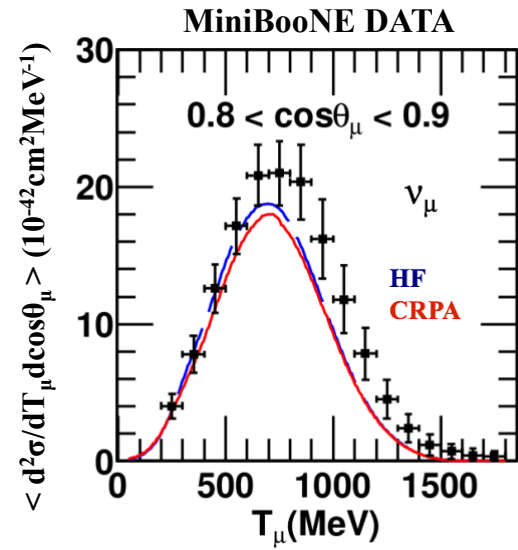


Comparison with (e,e') data for ^{12}C , ^{16}O , ^{40}Ca

VP, N. Jachowicz, PRC 92, 024606 (2015)



Comparison with neutrino data



Comparing RPA-based models

RPA polarization propagator: $\Pi = \Pi^0 + \Pi^0 V \Pi$

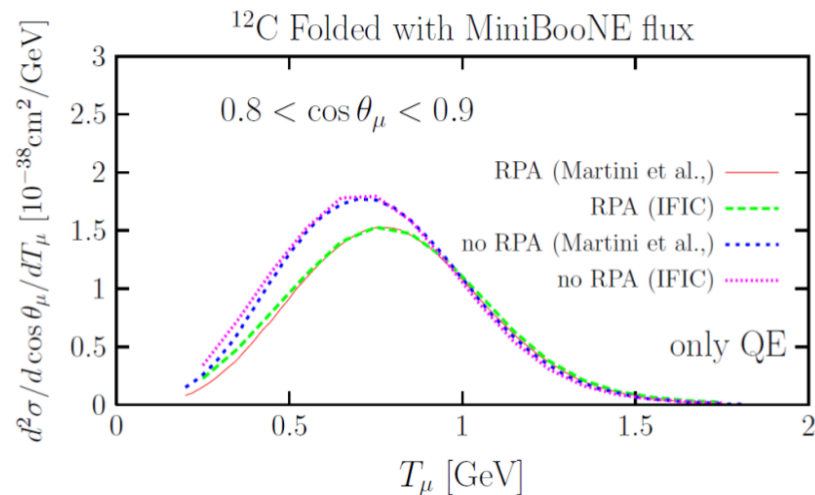
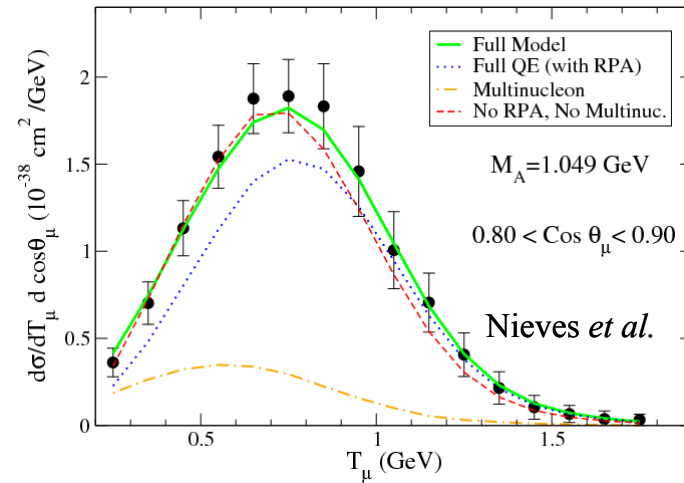
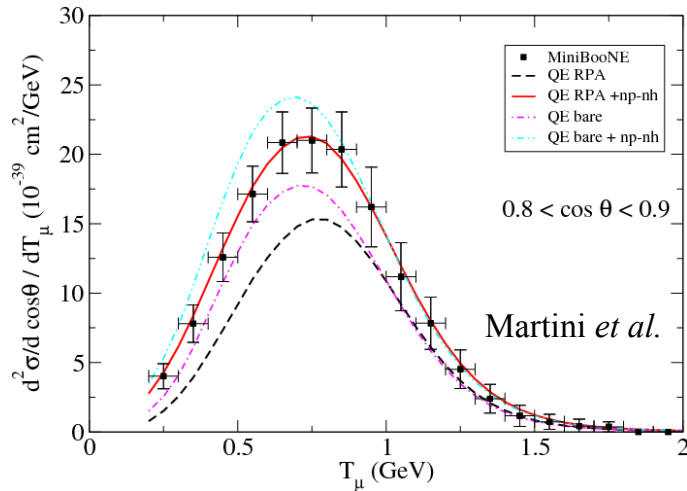
Comparing RPA-based models

RPA polarization propagator:

$$\Pi = \Pi^0 + \Pi^0 V \Pi$$

[Martini *et al.* and Nieves *et al.*]

Bare Propagator (RIFG) π exchange, ρ exchange, contact Landau-Migdal parameters



- Significant RPA quenching in both approaches.
- Genuine QE bare (RIFG) and RPA very similar in both approaches.

Comparing RPA-based models

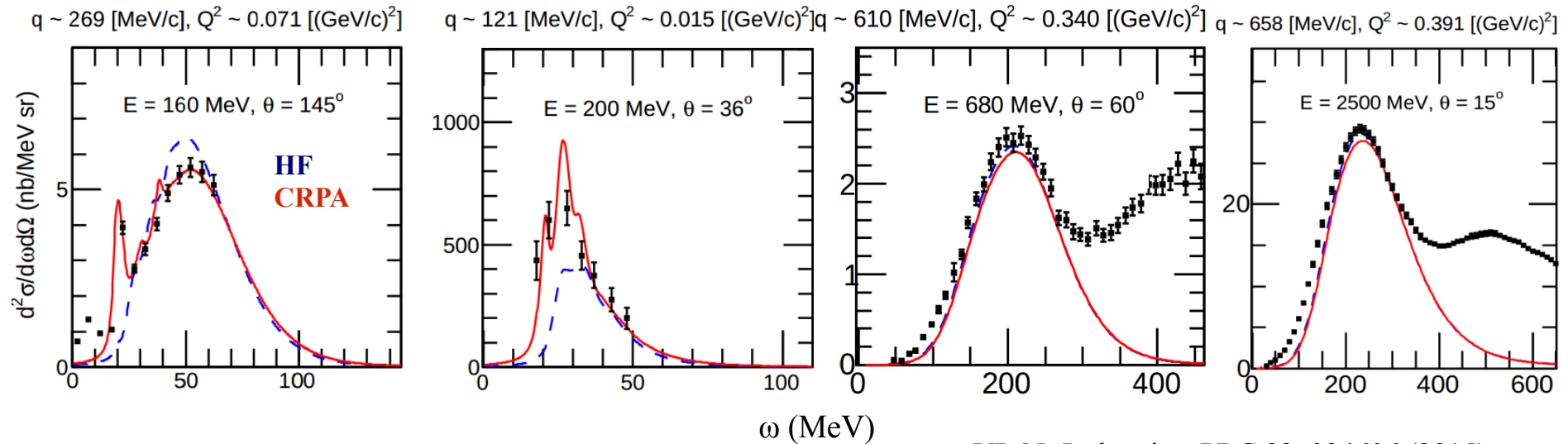
RPA polarization propagator:

$$\Pi = \Pi^0 + \Pi^0 V \Pi$$

[Pandey, Jachowicz *et al.*]

HF

Skyrme (SkE2)



VP, N. Jachowicz, PRC 92, 024606 (2015)

- At low ω , RPA (long-range correlations) describes the collective behavior of the nucleus (low-energy excitations).
- At high ω , RPA effects are smaller.
- Approach compares well with the (e,e') cross section.

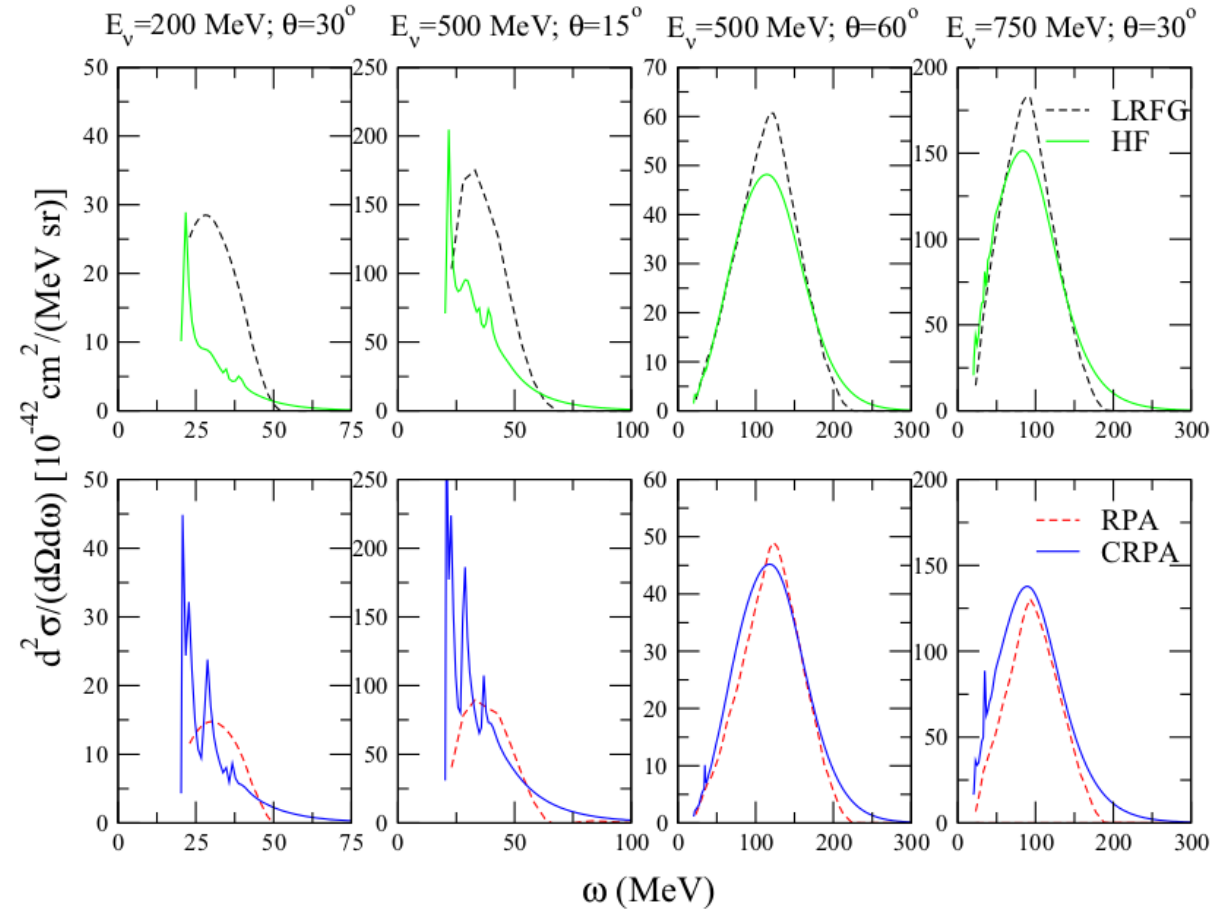
Comparing RPA-based models

For more details: M. Martini, N. Jachowicz, M. Ericson, and VP *et al.*, *PRC* 94, 015501 (2016)

LRFG, RPA: Martini, Ericson *et al.*

HF, CRPA: Pandey, Jachowicz *et al.*

- Important differences at both ends of the spectrum
 - Low-energy excitations at low ω
 - High ω tail



Comparing RPA-based models

Model	Starting point	N-N interaction	Shell effects	Low-energy excitations & Giant resonances	RPA effect
Martini, Ericson <i>et al.</i>	Local Fermi Gas	Meson -exchange (π, ρ, g')	No	No	Significant suppression (LLEE effect*)
Nieves <i>et al.</i>	Local Fermi Gas	Meson -exchange (π, ρ, g')	No	No	Significant suppression (LLEE effect*)
Pandey, Jachowicz <i>et al.</i>	Hartree-Fock	Skyrme	Yes	Yes	Describes low ω physics, not much effects at higher ω

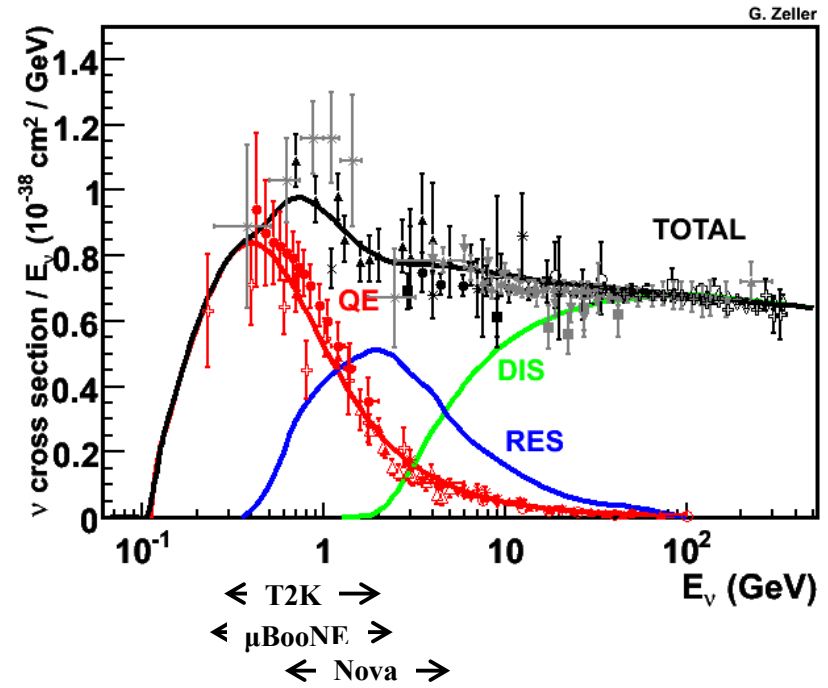
- Significant differences between RPA and CRPA approach, at both ends of the (one-body) ω spectrum.

*Lorentz-Lorentz-Ericson-Ericson effect: accounts for the possibility of a Δ -hole excitation in the RPA chain

Impact of low-energy excitations

What neutrino brings: E_ν

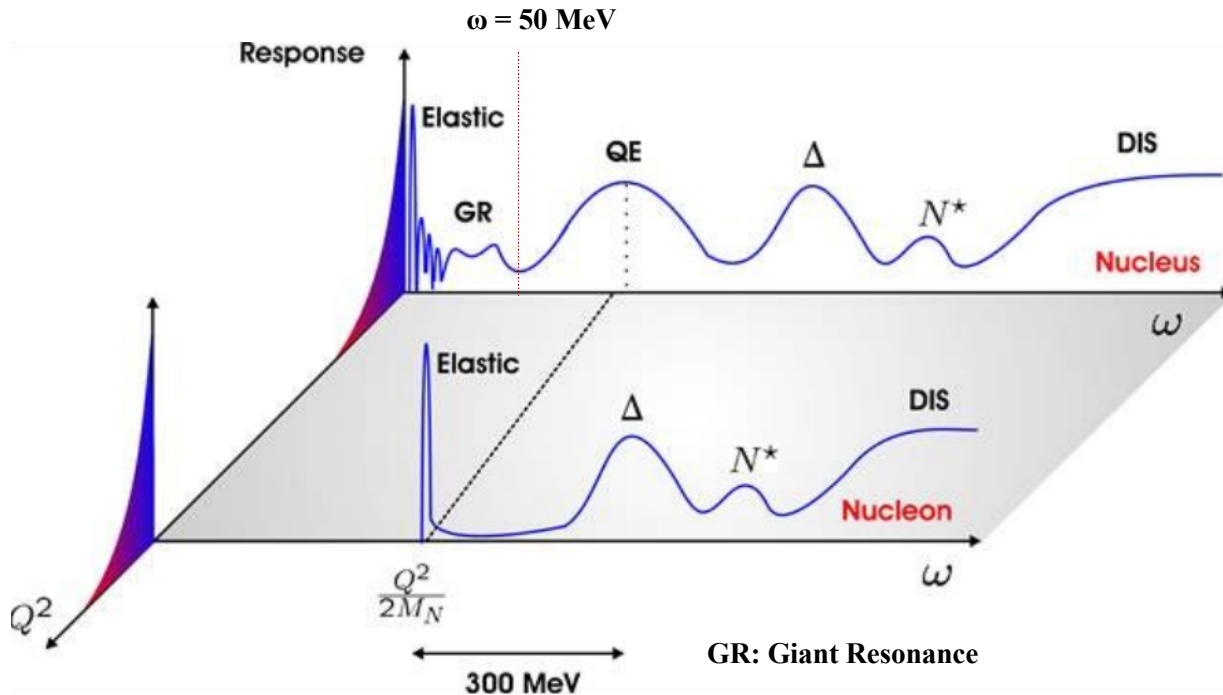
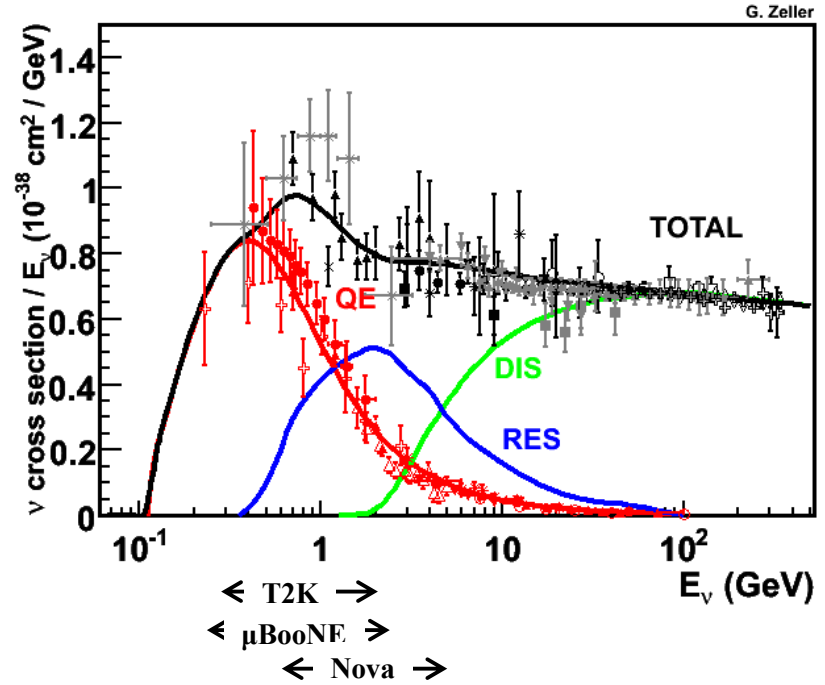
Cross section (integrated over full phase-space)
in terms of incoming neutrino energy E_ν .



Impact of low-energy excitations

What neutrino brings: E_ν

Cross section (integrated over full phase-space) in terms of incoming neutrino energy E_ν .



What nucleus cares: ω

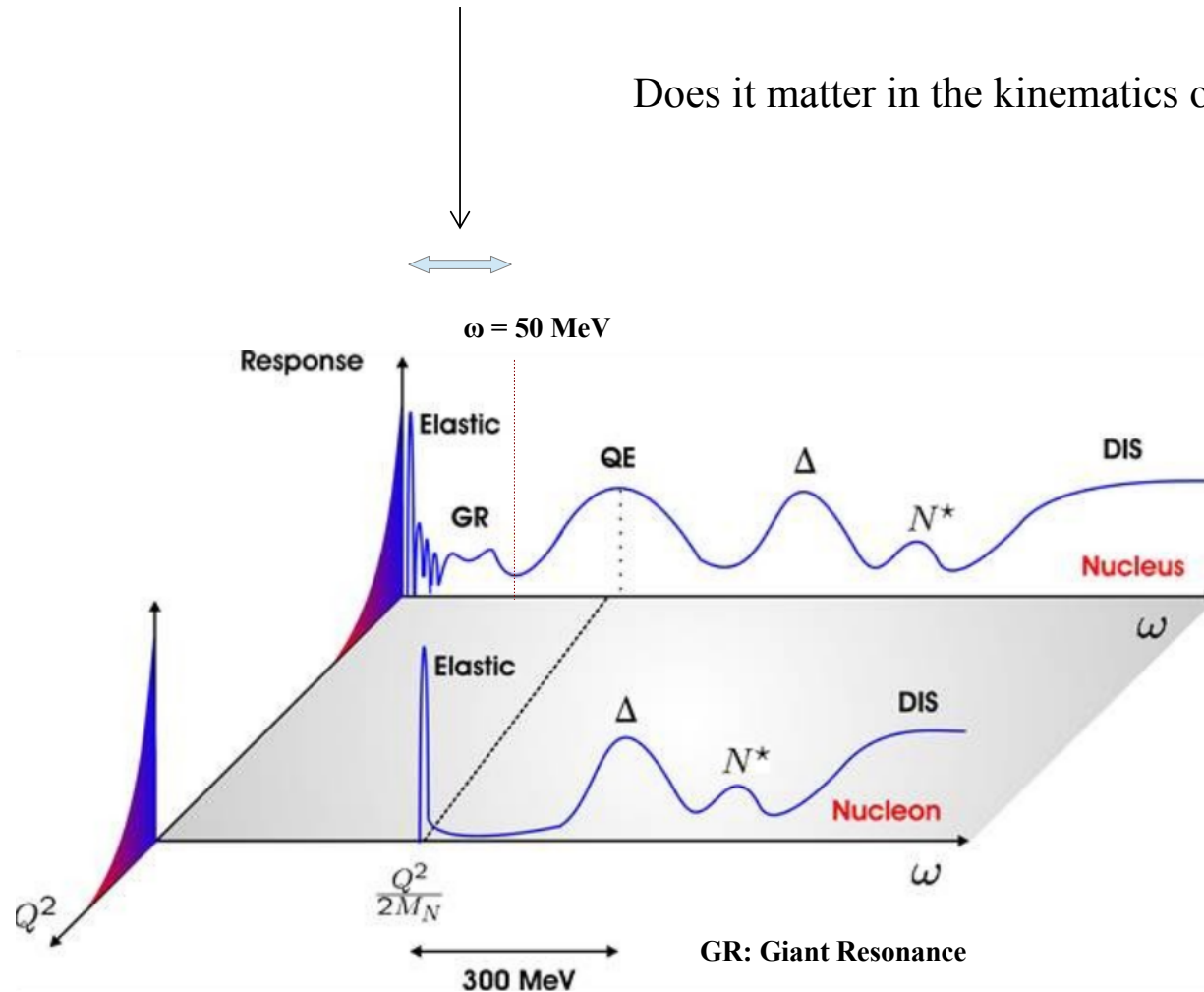
i.e. how much energy is transferred to the nucleus (ω).

For a given E_ν , it depends on other factors, such as on lepton scattering angle.

Impact of low-energy excitations

What is missed in the translation.

Does it matter in the kinematics of our interest?



What nucleus cares: ω

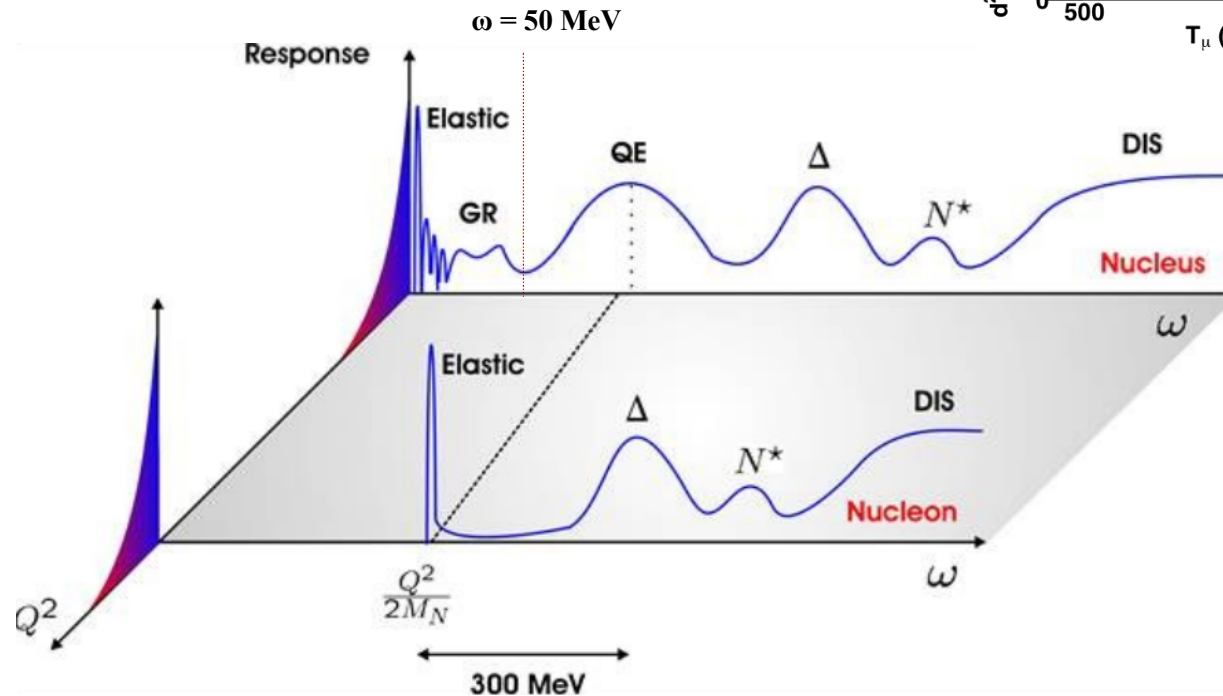
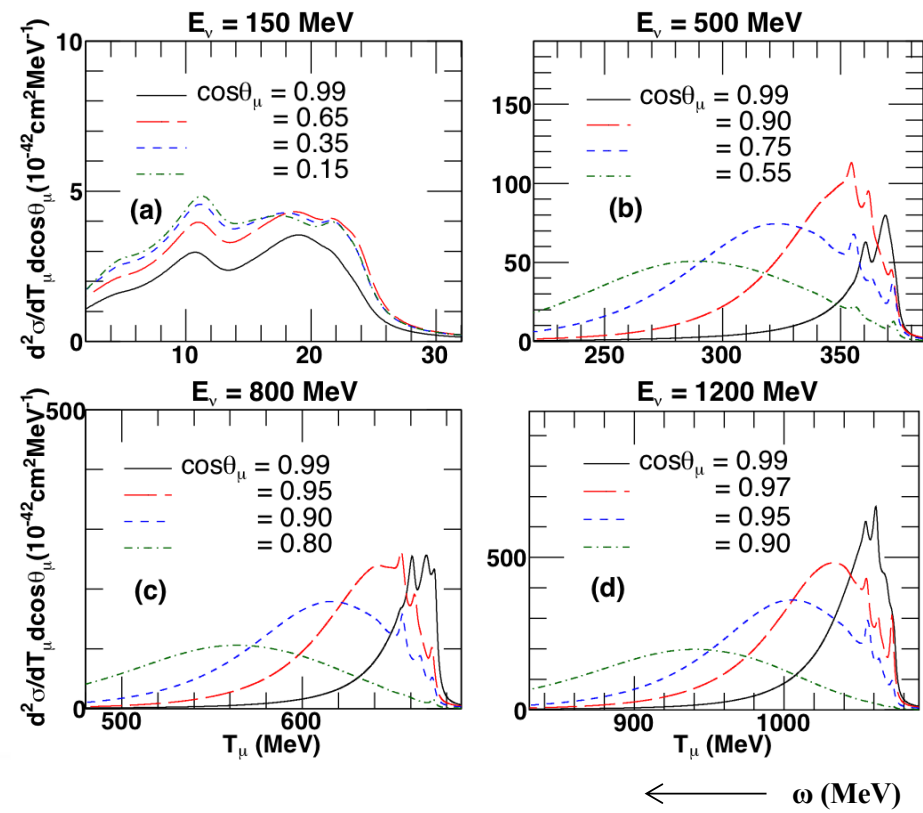
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Impact of low-energy excitations

Low E_ν : cross section is dominated by low-energy excitations.

E_ν at the peak of MicroBooNE/T2K, forward scattering receive contribution from low-energy excitations.



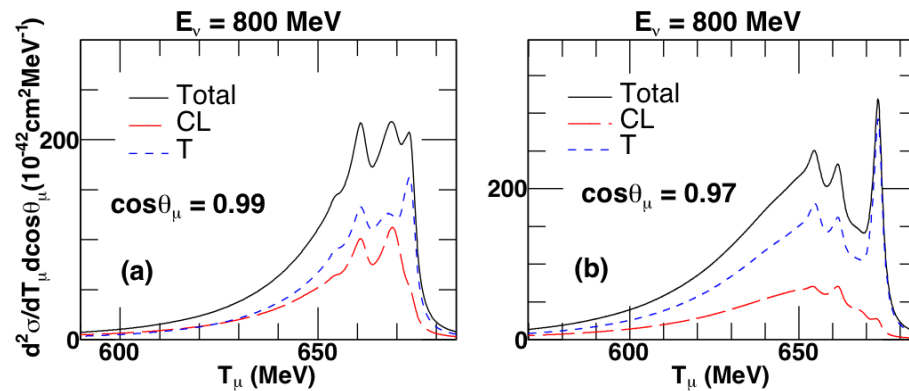
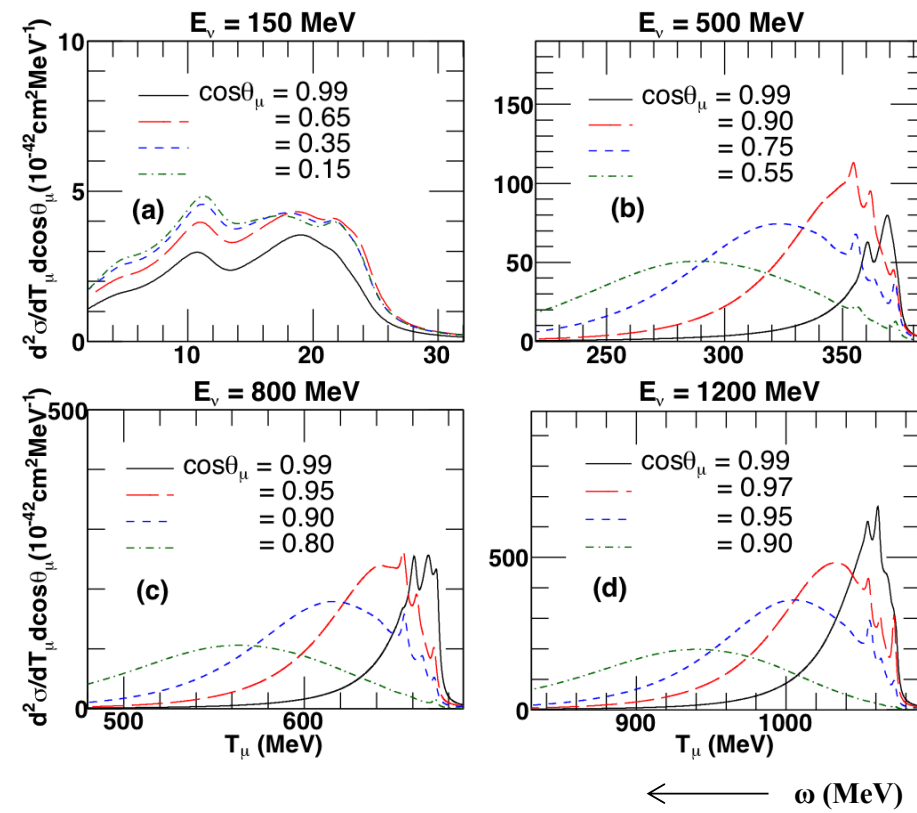
VP, N. Jachowicz et al, PRC 92, 024606 (2015)

Impact of low-energy excitations

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E_ν at the peak of T2K flux, forward scattering receive contribution from low-energy excitations.

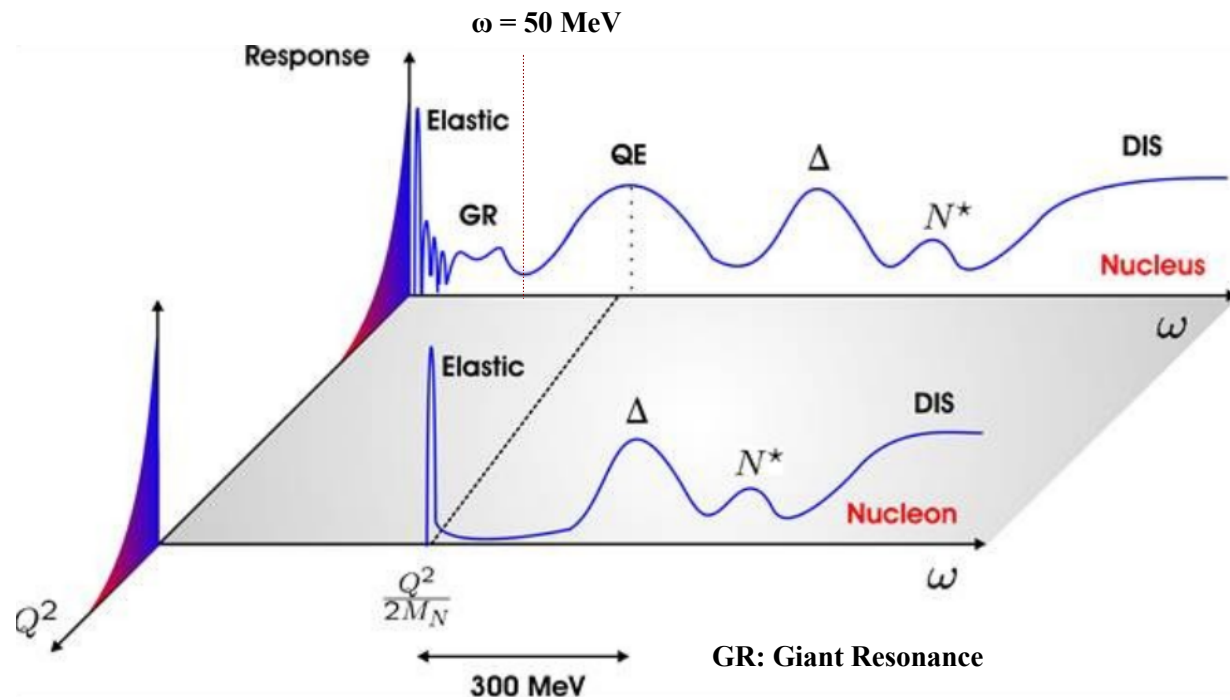
- The forward we go in scattering angle, longitudinal contribution starts competing with the transverse one (at intermediate energy).
- At low-energies and forward scattering, longitudinal response dominates over transverse one.

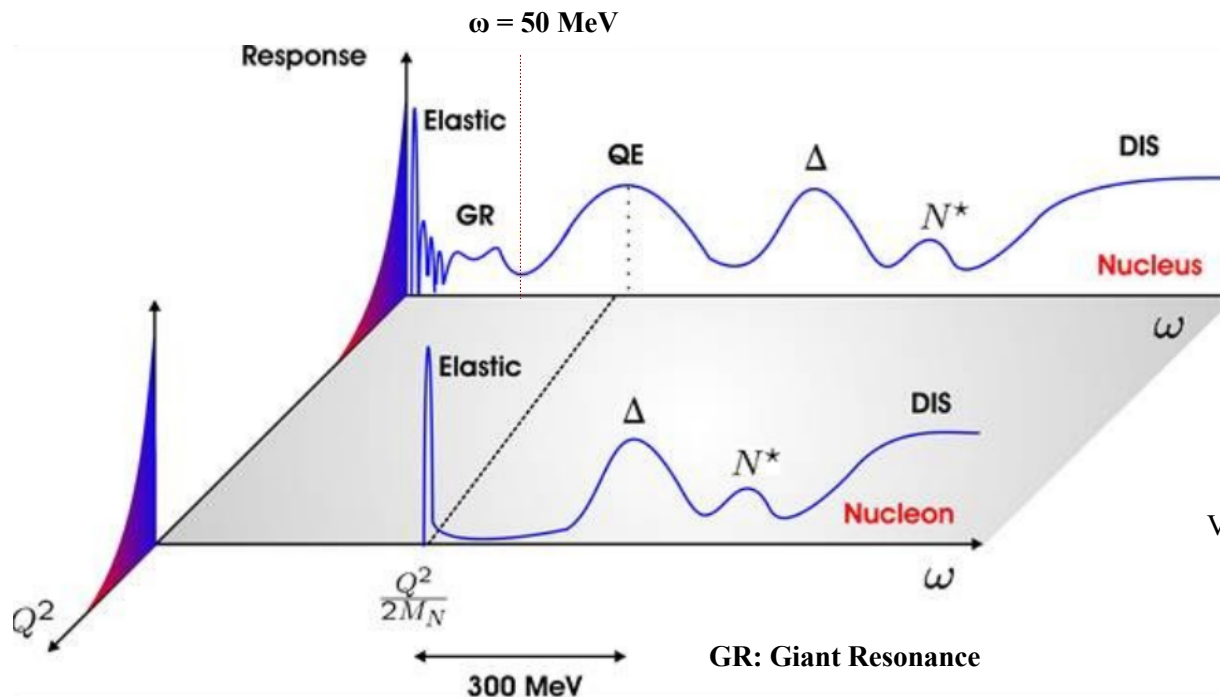
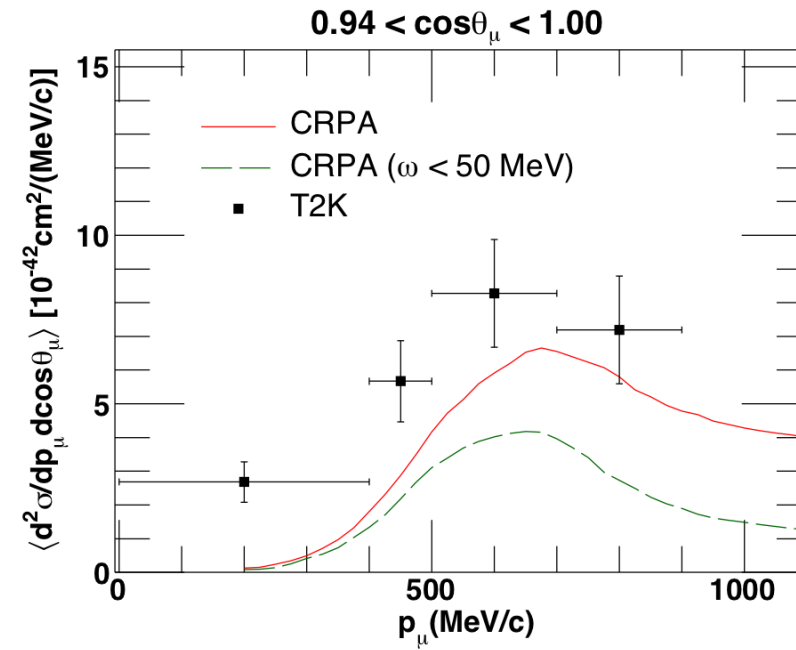
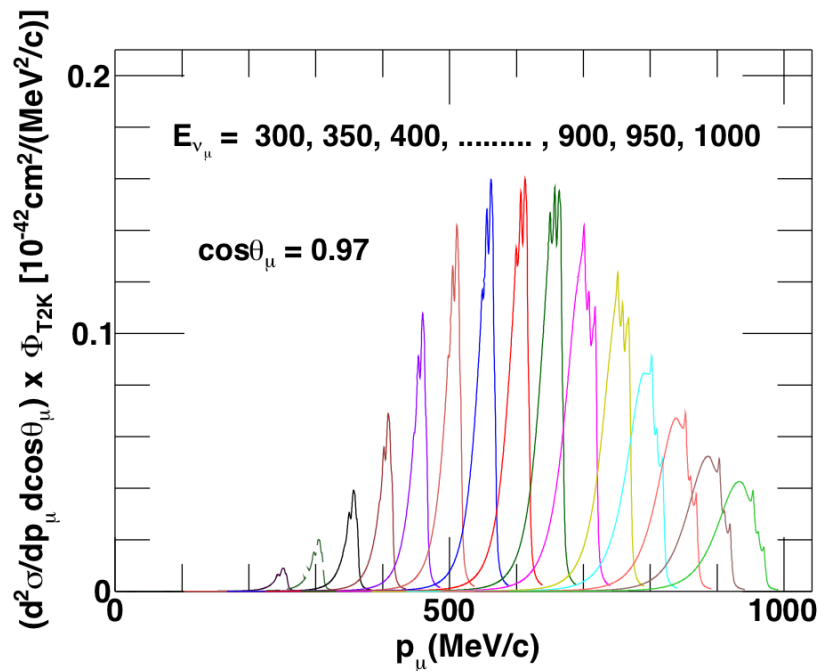


VP, N. Jachowicz et al, PRC 92, 024606 (2015)

Impact of low-energy excitations

Does it affect the flux-folded cross section?
And how much?





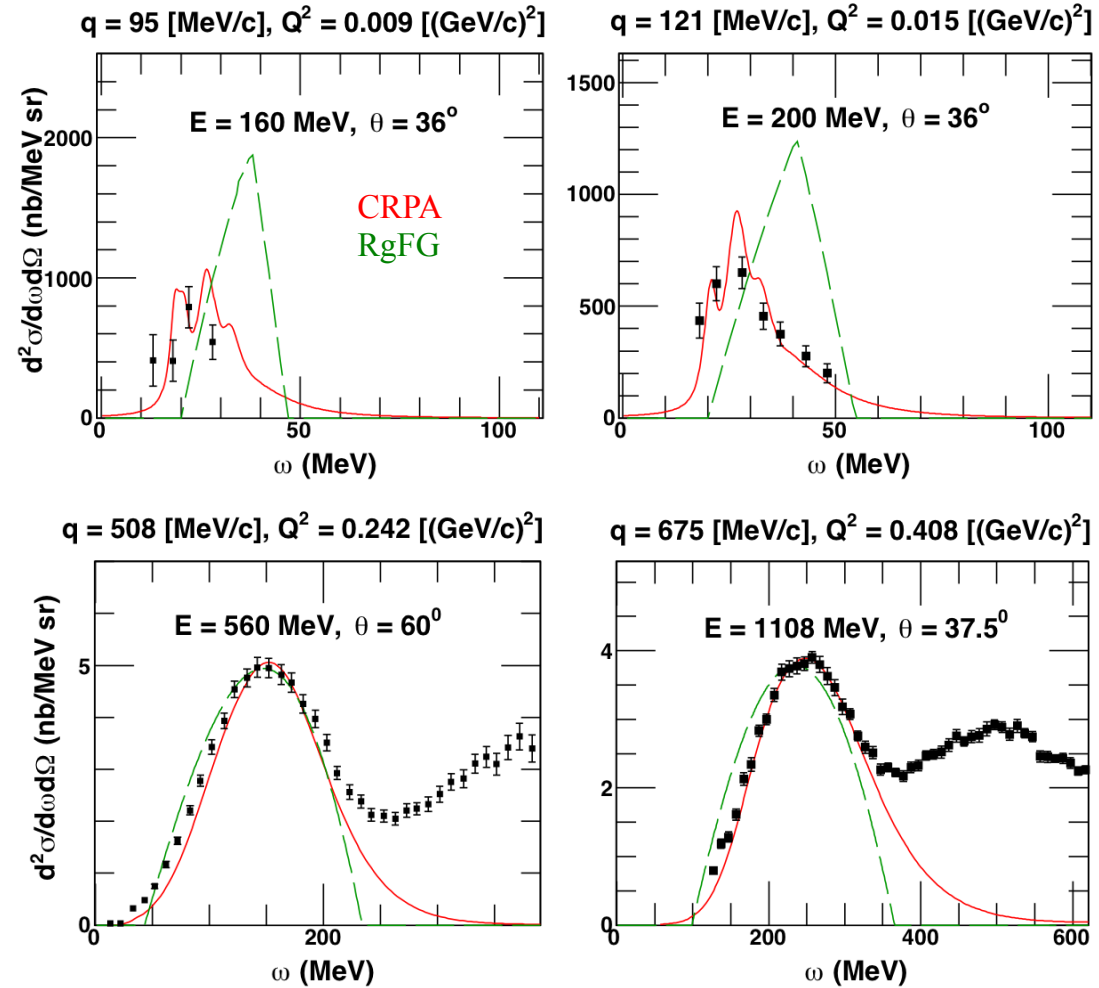
~ 50% of the flux folded cross section in this forward bin emerges from low-energy nuclear excitations.

VP, N. Jachowicz, Phys.Rev. C94, 054609 (2016).

Impact of low-energy excitations

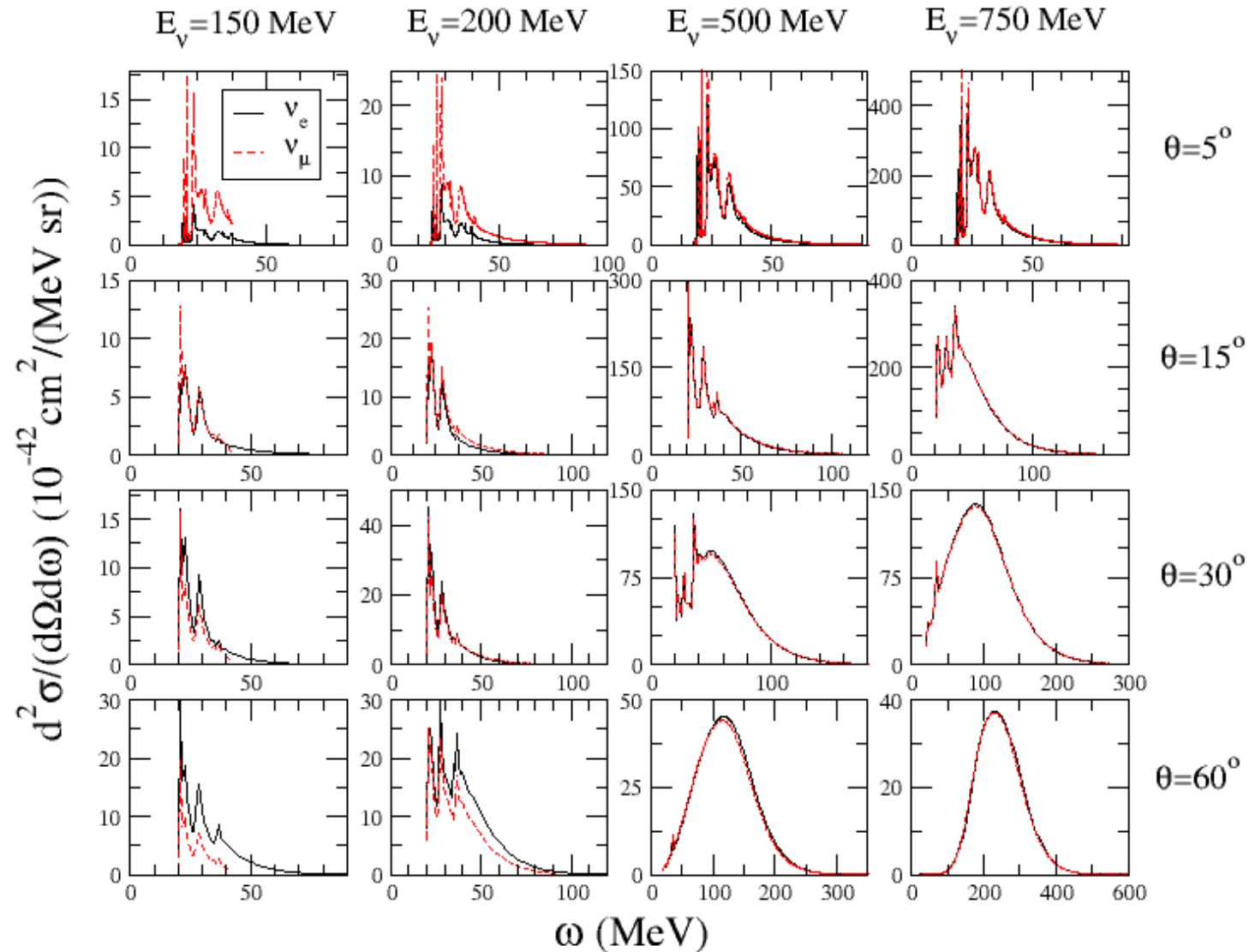
How does RgFG model describes low-energy excitations – it does not!

Example from ^{12}C (e, e')



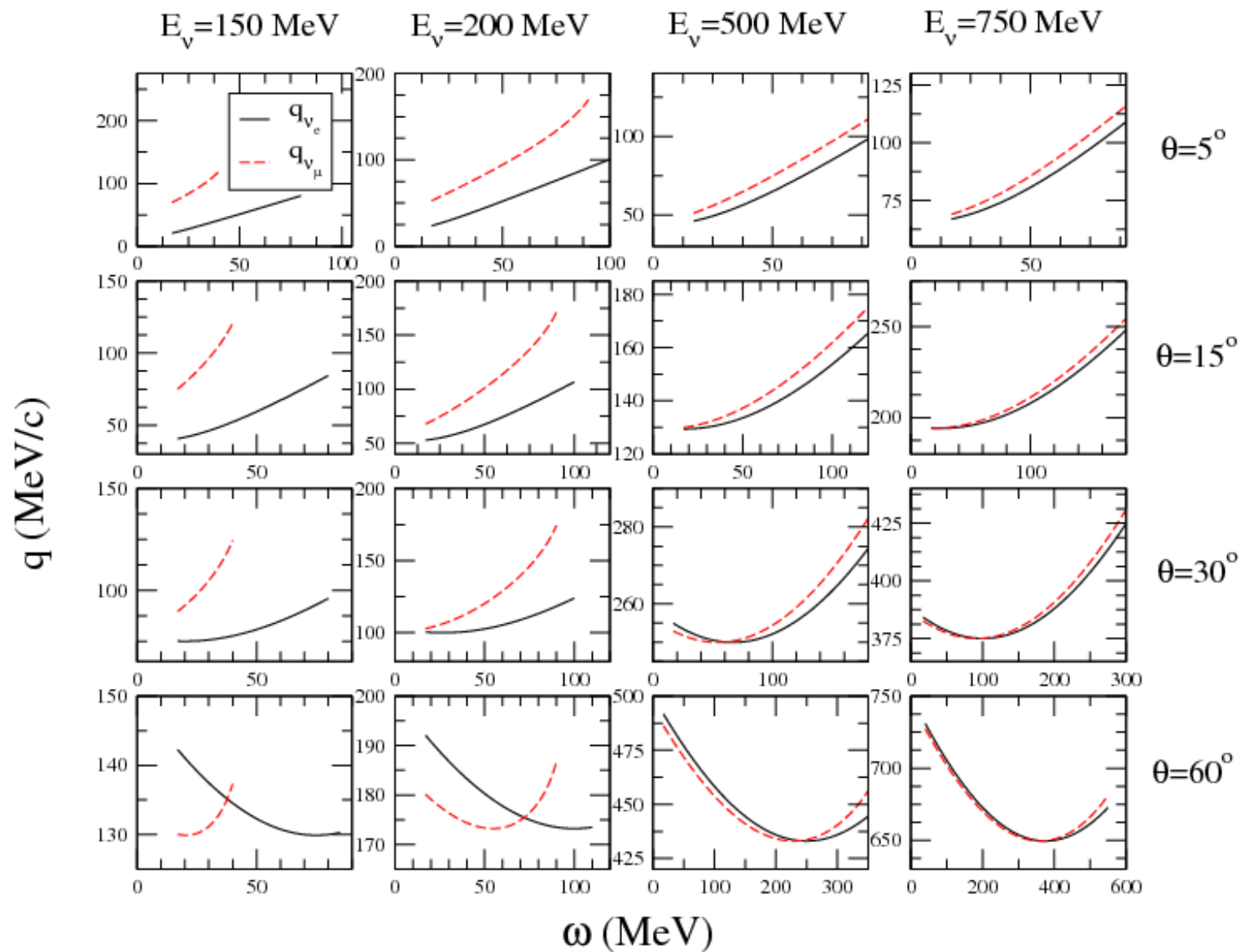
Impact of low-energy excitations

The effect of low-energy nuclear excitations on ν_e vs ν_μ cross section



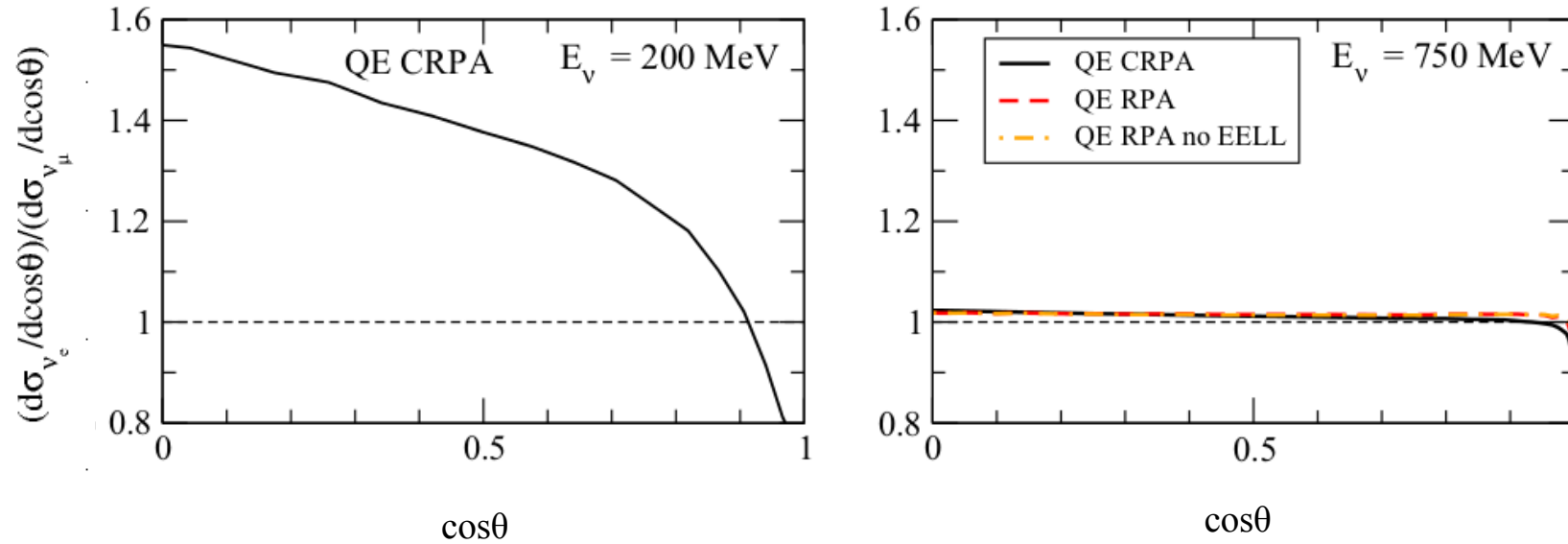
Impact of low-energy excitations

The effect of low-energy nuclear excitations on ν_e vs ν_μ cross section



Impact of low-energy excitations

The effect of low-energy nuclear excitations on ν_e vs ν_μ cross section



Low-energy nuclear excitations are vital

- At low E_ν
- At intermediate E_ν and forward scattering
- Differentiating between ν_e and ν_μ cross section (at low E_ν)

Summary

The Impact of nuclear effects on accelerator-based neutrino-oscillation physics



- Accelerator-based neutrino-oscillation program, for the precision measurement of neutrino oscillation parameters and hopefully establishing CP violation in leptonic sector, is moving with full steam.



- There are still major issues related to identification of basic processes contributing to the neutrino-nucleus signal in a detector and the reduction of systematic uncertainties.

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- The Monte-Carlo generators are still based on inadequate models and implementing new models into current generators is complex and slow.

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- Next generation experiments (MicroBooNE, ICARUS, SBND, and DUNE) employ LArTPC technology but a little (almost nothing) is known about Ar nucleus and its electroweak response.



- A new Ar(e,e'p) experiment at Jlab (was recently approved and now also already finished collecting data) will provide spectroscopic factors and cross sections on Ar.