M	EOS	ν_{w_1} (Hz)	τ_{w_1} (μs)	$\nu_{w_1^{II}}$ (Hz)	$\tau_{w_1^{II}}$ (μs)
	APR2	8437	24.6	2081	12.3
	APRB200	8496	24.6	2162	12.2
	APRB120	8496	24.6	2162	12.2
	BBS1	7989	24.5	1580	12.8
	BBS2	8925	24.4	2545	11.8
	G240	7442	24.6	1019	13.5
	SS1	9990	26.1	4498	10.9
	SS2	9997	26.1	4520	10.8
1.4 $M_\odot$	APR2	8156	29.5	3040	13.1
	APRB200	8190	29.5	3110	13.1
	APRB120	8189	29.6	3110	13.1
	BBS1	7785	29.0	2495	13.6
	G240	7588	28.6	2157	13.9
	SS1	9655	34.5	6050	11.9
	SS2	–	–	–	–
1.5 $M_\odot$	APR2	8016	32.4	3499	13.6
	APRB200	8040	32.5	3559	13.5
	APRB120	8042	32.5	3562	13.5
	BBS1	7691	31.7	2947	14.0
	G240	7910	31.7	3105	13.6
1.8 $M_\odot$	APR2	7590	45.2	4865	15.2
	APRB200	7593	43.3	4894	15.2
	APRB120	7683	46.3	5096	15.1
	BBS1	7436	43.3	4399	15.3
2 $M_\odot$	APR2	–	–	–	–
	APRB200	–	–	–	–
	BBS1	–	–	–	–
M_{max}	APR2	–	–	–	–
	APRB200	–	–	–	–
	APRB120	–	–	–	–
	BBS1	–	–	–	–
	BBS2	–	–	–	–
	BBS2	9567	25.3	3436	11.4
	G240	8616	36.2	4562	13.3
	SS1	–	–	–	–
	SS2	–	–	–	–

TABLE I: The frequencies and the damping times of the *axial* QNM for $l = 2$ are tabulated for assigned values of the stellar mass and for the considered EOS's. As in Table III, when the mode frequency cannot be computed the slot is left empty.