

March 31

## Chiral Hamiltonians

Start from an effective Lagrangian involving pions and nucleons, constrained by the broken chiral symmetry of strong interaction

Infinite number of terms

$$\mathcal{L} = \mathcal{L}^{(0)} + \left(\frac{Q}{\Lambda_b}\right)^2 \mathcal{L}^{(2)} + \left(\frac{Q}{\Lambda_b}\right)^4 \mathcal{L}^{(4)} + \dots$$

$\Lambda_b \sim 800 \text{ MeV}$

# Main advantages:

- systematic scheme
- two- and many-nucleon forces derived within

a consistent framework

# Drawbacks:

- Cutoff dependence
- Momentum expansion

In degenerate neutron matter

$$k_F = (3\pi^2 \rho_B)^{1/3}$$

neutron-neutron collisions can only involve momenta  $\sim k_F$  (Pauli blocking)  $\Rightarrow$  energy  $\sim$

$$E_{lab} = 2 E_{cm}$$

$$= \frac{2}{m} (3\pi^2 \rho_B)^{2/3}$$

ability to describe  
phase shifts  $\Leftrightarrow$   
ability to describe  
dense matter.

(see figure 5)

Conclusion: chiral interactions provide a viable alternative to purely phenomenological Hamiltonians, but cannot reliably used @ densities larger than  $1-2\rho_0$   
( $\rho_0 = 0.16 \text{ fm}^{-3} \approx 2.7 \cdot 10^{14} \text{ g cm}^{-3}$ )

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Many-body theory

$$H = T + V + U$$

Even the simplest problem: solving

$$H |\Psi_0\rangle = E_0 |\Psi_0\rangle$$

involve prohibitive  
computational difficulties  
for  $A > 4$ .

Structure of  $|\Psi_0\rangle$  involves  
strong spin-isospin

dependence

# of spin isospin  
configurations

$${}^4\text{He} \quad \sim \quad 10^2$$

$${}^{12}\text{C} \quad \sim \quad 10^6$$

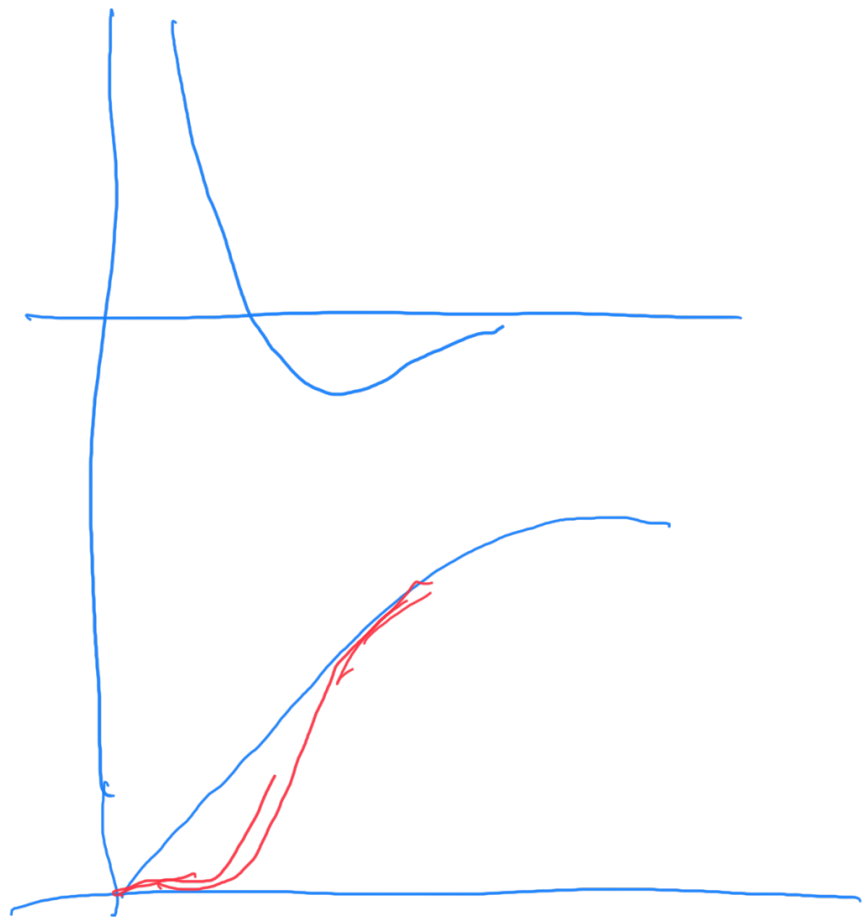
- Brute force solution  
using stochastic methods
  - Approximate solutions
    - perturbation theory  
made difficult  
by strong range repulsion
- Consider a nucleon  
pair in uniform  
matter

$$\psi_0 \sim \frac{\sum (kr)}{kr}$$

First-order perturbation theory

$$\Delta E \sim \langle \psi_0 | V | \psi_0 \rangle$$

large



Convergence  $\rightarrow$  converges  $\psi_0$

$\psi$ -method - Three-nucleon forces  
Variational  $\rightarrow$  accuracy.

Quantum Monte Carlo

Imaginary time Schrödinger equation

$$-\frac{\partial \psi_0}{\partial \tau} = H \psi_0$$

$$\psi(x, \tau) = \int dx' G(x, x', \tau) \psi(x', \tau)$$

$$G(x, x', \tau) = \langle x | e^{-H\tau} | x' \rangle$$

Trial wave function

$$|\tilde{\psi}_0\rangle = \sum_n c_n |\psi_n\rangle$$

$$|\psi(0)\rangle = |\tilde{\psi}_0\rangle$$

$$\lim_{\tau \rightarrow \infty} |\psi(\tau)\rangle = |\psi_0\rangle$$

$$\begin{aligned} & \tau \rightarrow \infty \quad | \psi \rangle \langle \psi | - (E_0 - t_{ul}) \tau \\ & = \lim_{\tau \rightarrow \infty} e^{-\tau (E_0 - t_{ul})} \quad @_0 | \psi_0 \rangle \end{aligned}$$

QMC possible for  
 pure neutron matter (64  
 neutrons in a box)

<sup>150 spin up</sup>  
 - Symplectic matter still  
 involves uncertainties.