

# Relativity

(A) Boost corrections

(B) QFT + RMF approximation

$$\begin{array}{c} \xrightarrow{\quad} \vec{P}_i + \vec{P}_j \\ \text{(A)} \quad \sigma_{ij} \rightarrow \sigma_{ij}(\vec{P}) = \sigma_{ij} + \delta\sigma(\vec{P}) \end{array}$$

$$\delta\sigma_{ij}(\vec{P}) = \frac{\vec{P}^2}{8m} \sigma_{ij}$$

$$+ \frac{1}{2} [(\vec{P} \cdot \vec{r}_{ij}) (\vec{P} \cdot \vec{\nabla}_{ij}), \sigma_{ij}]$$

$$+ \frac{1}{2} [(\vec{\sigma}_i - \vec{\sigma}_j) \times \vec{P} \cdot \vec{r}_{ij}, \sigma_{ij}]$$

QFT? needed when number of particles is not conserved

replace wave function with field

starting point is a lagrangian density (similar to Yukawa's)

simplest implementation: Weicko model (mid 1970s)

degrees of freedom

- nucleon  $\Psi_N = \begin{pmatrix} \uparrow \\ \downarrow \end{pmatrix}$

- vector-isoscalar field  $\omega^\mu$   $m_\omega \sim 780$  MeV short range repulsive

- scalar-isoscalar field  $\sigma$   $m_\sigma \sim 550$  MeV (not a physical, that is, observed particle)

$$\mathcal{L}_0 = \bar{\Psi} (i \not{\partial} - m) \Psi + \frac{1}{2} \partial_\mu \sigma \partial^\mu \sigma - \frac{1}{2} m^2 \sigma^2 - \frac{1}{4} F_{\mu\nu} F^{\mu\nu} + \frac{1}{2} m_\omega^2 \omega_\mu \omega^\mu$$

$$F_{\mu\nu} = \partial_\nu \omega_\mu - \partial_\mu \omega_\nu$$

$$\mathcal{L}_{int} = g_\sigma \bar{\Psi} \Psi \sigma - g_\omega \bar{\Psi} \gamma^\mu \Psi \omega_\mu$$

Eq. of motion cannot be solved using perturbation theory

$$(\square + m_\sigma^2) \sigma = g_\sigma \bar{\Psi} \Psi$$

$$(\square + m_\omega^2) \omega_\mu - \partial_\mu (\partial_\nu \omega^\nu) = g_\omega \bar{\Psi} \gamma_\mu \Psi$$

$$\Gamma \sim \langle \bar{\Psi} \Psi \rangle \sim \langle \omega^\mu \omega_\mu \rangle \sim \langle (\square - g_\sigma \sigma) \Gamma \rangle = 0$$

$$\int \sigma_\mu (i\partial - \gamma \omega \gamma) \psi$$

MFA  $\sigma \rightarrow \langle \sigma \rangle$

$$\omega_\mu \rightarrow \langle \omega_\mu \rangle \rightarrow \langle \omega_0 \rangle$$

( $\langle \omega_i \rangle = 0$ )

for example

$$m_\sigma^2 \langle \sigma \rangle = g_\sigma \langle \bar{\psi} \psi \rangle$$

$\langle \rangle =$  sum over all states  
belonging to the  
Fermi sea

$$\left[ (i\partial - g_\omega \gamma_0 \langle \omega_0 \rangle) - (m - g_\sigma \langle \sigma \rangle) \right]$$

$$\mathcal{L}_{\text{MFA}} = \bar{\psi} \left[ \right] \psi - \frac{1}{2} m_\sigma^2 \langle \sigma \rangle^2 + \frac{1}{2} m_\omega^2 \langle \omega_0 \rangle^2$$

$\psi = 0$

Eq. of motion from energy-momentum tensor

$$T^{\mu\nu} = \frac{\partial \mathcal{L}}{\partial \partial_\mu \psi} \partial^\nu \psi + \frac{\partial \mathcal{L}}{\partial \partial_\mu \bar{\psi}} \partial^\nu \bar{\psi} - g^{\mu\nu} \mathcal{L}$$

$$\Rightarrow i \bar{\Psi} \gamma^\mu \partial^\nu \Psi - g^{\mu\nu} \mathcal{L}$$

EOS from

$$\epsilon = T^{00}$$

$$P = \frac{1}{3} T^{ii}$$

at fixed baryon density  $\langle \bar{\Psi} \Psi \rangle$

$$\Rightarrow P = P(\epsilon)$$