

Relativistic Mean field,
continued

$$(\square + m_\sigma^2)\sigma = g_\sigma \bar{\Psi} \Psi$$

$$(\square + m_\omega^2)\omega_\mu + \partial_\mu(\partial^\nu \omega_\nu) = g_\omega \bar{\Psi} \gamma_\mu \Psi$$

$$[\gamma_\mu (i\partial^\mu - g_\omega \omega^\mu) - (m - g_\sigma \sigma)]\Psi = 0$$

mean-field approximation

$$\sigma \rightarrow \langle \sigma \rangle = \text{constant}$$

$$\omega_\mu \rightarrow \langle \omega_\mu \rangle \begin{cases} 0 & \mu = 1, 2, 3 \\ \text{constant} & \mu = 0 \end{cases}$$

$$\begin{array}{l} \text{effective } \left\{ \begin{array}{l} \text{momentum} \\ \text{mass} \end{array} \right. \end{array} \quad \begin{array}{l} k^\mu - g_\omega \omega^\mu \\ m^* = m - g_\sigma \sigma \end{array}$$

EOS from energy-momentum tensor

$$T^{\mu\nu} = \frac{\partial \mathcal{L}}{\partial \partial_\mu \psi} \partial^\nu \psi + \frac{\partial \mathcal{L}}{\partial \partial_\mu \bar{\psi}} \partial^\nu \bar{\psi} - g^{\mu\nu} \mathcal{L}$$

$$T^{\infty} = \epsilon$$

$$\frac{1}{3} T^{ii} = P$$

③ fixed baryon density

Result only depends on

$$\frac{g_\sigma}{m_\sigma} \quad \frac{g_\omega}{m_\omega}$$

model fully determined from equilibrium properties of isospin-symmetric matter

